

# Some results of comparison between the lower thermosphere zonal winds as seen by the ground-based radars and WINDII on UARS

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The seasonal variations of the zonal winds measured by meteor/MF radars and by the wind-imaging interferometer (WINDII) on board of the UARS satellite, are analyzed and compared with the ground-based Global Empirical Wind Model (GEWM) and satellite-based WINDII prevailing wind model, which are independently derived from the observational data sets. A general consistency in the main global scale wind structures is observed. The seasonal variations of zonal wind in the both models are described by almost the same annual and semi-annual components. But systematic bias is found for the annual mean zonal winds of the two models, with the GEWM winds generally smaller than those of WINDII by factor 2–2.5. This bias is practically independent of altitude and can be described by a term of  $A \cos 4x$ , where  $A$  is about 20 m/s and  $x$  is colatitude. Possible origin of this offset is discussed.

## 1. Introduction

Various global empirical wind models play an important role in our understanding of main regularities in the upper mesosphere/lower thermosphere dynamics. Among others, the models derived from the ground-based radar wind measurements have been published by Portnyagin (1984, 1987), Portnyagin and Solovjova (1992). Constant accumulation of data sets collected from ground- and space-based instruments has continuously stimulated the development and improvement of these models. The updated version of the models (hereafter called as GEWM) is based on the most recently available ground-based meteor radar and MF (partial reflection) wind measurements. Meanwhile, direct wind observations from UARS experiment provide unique global data sets with a substantial increase in the quantity of available wind measurements in the mesosphere and lower thermosphere. Based on the WINDII green line wind data, Wang *et al.* (1997) have proposed an empirical horizontal wind model. Here we present the results of comparison between zonal winds retrieved from the GEWM and the WINDII model to show their general consistency and significant difference.

## 2. Global Empirical Wind Model

The discussed GEWM is constructed by assimilating the all available data from meteor radar and MF wind measurements in the upper mesosphere/lower thermosphere. The measurement's sites and the observational periods are presented in Table 1. The meteor radar wind data (including those obtained for the average height of about 95 km) are seen to give us the bulk ground-based information about the wind regime in the lower thermosphere.

The first step in construction of the model is to interpolate (and somewhere extrapolate) the all height-resolved experimental monthly mean wind values (height profiles) to constant height levels with a standard step (usually 1 km). Then the wind values at each height level were interpolated over latitude using a routine cubic spline procedure. The obtained values were additionally smoothed over latitude with help of the Legendre function's decomposition. This procedure is correct for the zonal mean wind models. (Actually, the zero winds at the poles are assumed.)

The next step is adaptation of the above-derived preliminary model to the meteor wind data without height resolution. These data are usually related to an average height of about 95 km at different latitudes, and show very consistent and regular seasonal behaviour, practically the same as that deduced from the meteor radar wind measurements with height resolution. We have estimated the effect of averaging over height on the monthly mean winds by using the measurements of the Kazan meteor radar with height resolution Lysenko *et al.* (1994). The general conclusion is that the seasonal course of the height-averaged and height-resolved monthly mean meteor wind data are well matched at the particular height of about 95 km, and that climatic features for their seasonal variations are very persistent, independent of longitude and observational periods and consistent for all types of devices. This conclusion is supported by comparison of the wind data measured at other latitudinal belts with and without height resolution (see Table 1).

Based on these results, we tuned the preliminary wind profiles at about 95 km to include the height-averaged data observed at the particular sites. Then, the smoothing procedure over latitude, as described at the first step, were repeated for the whole set of the wind profiles. In the two steps of constructing the model, the statistical weight (in a climatological sense) of the measurement data was also taken into account.

Table 1. Data base.

Station	Location	Method	Observing period	References
Heiss I.	80.5°N, 58°E	MR*	1.1965–10.1985	Portnyagin (1986)
Tromso	70°N, 19°E	MF	1987–1989	Manson <i>et al.</i> (1991)
Kiruna**	68°N, 20°E	MR*	1974–1975	Manson <i>et al.</i> (1985)
Poker Flat	65°N, 147°W	MR	7.1980–12.1984	Manson <i>et al.</i> (1987)
College	65°N, 148°W	MR*	1.1967–8.1968	Hook (1970)
Tomsk	57°N, 85°E	MR*	10.1965–12.1966	Nazarenko (1968)
Kazan	56°N, 49°E	MR	1986–1988	Fakhrutdinova (1991), Sidorov <i>et al.</i> (1988)
Obninsk	55°N, 37°E	MR*	1964–1995	Portnyagin (1986)
Kuhlungsborn	54°N, 12°E	MR*	1977–1980 1–3.1990	HHI Geophys. Data, 1977–1980 Singer <i>et al.</i> (1994)
Juliusruh	54.6°N, 13.5°E	MF	1990–1991	Schminder <i>et al.</i> (1994)
Jodrell Bank	53°N, 2°W	MR*	1953–1958	Greenhow and Neufeld (1961)
Saskatoon	52°N, 107°W	MF	1979–1982	Manson <i>et al.</i> (1985)
Sheffield	53.3°N, 3.8°W	MR*	1–3.1990	Singer <i>et al.</i> (1994)
Badary	52°N, 102°E	LF*	1975–1981	Petruchin (1983)
Collm	52°N, 15°E	LF	1983–1986 1990–1991	Schminder and Kurschner (1988) Schminder <i>et al.</i> (1994)
Kharkov	50°N, 36°E	MR* MR	1980–1983 1987	Kalchenko (1987) Kascheev <i>et al.</i> (1988)
Kiev	50°N, 31°E	MR*	9.1964–2.1966	Lysenko <i>et al.</i> (1969)
Khabarovsk	49°N, 135°E	MR*	1976–1985	Makarov (1988)
Volgograd	49°N, 44°E	MR*	1978–1985	Portnyagin (1986)
Garchy**	47°N, 3°E	MR	1970–1976	Manson <i>et al.</i> (1985)
Monpazier**	45°N, 1°E	MR	1975–1980	Manson <i>et al.</i> (1985)
Bologna	45°N, 12°E	MR*	1–3.1990	Singer <i>et al.</i> (1994)
Durham	43°N, 71°W	MR	1978/79/84	Manson <i>et al.</i> (1987)
Frunze	43°N, 73°E	MR*	1964–1982	Karimov (1984)
Yambol	42.5°N, 26.5°E	MR*	2–3.1987	Lysenko <i>et al.</i> (1988)
Urbana	40°N, 88°W	MR	1991–1992	Franke <i>et al.</i> (1993)
Dushanbe	38°N, 68°E	MR*	1968–1969	Babadjanov <i>et al.</i> (1974)
Ashkhabad	37°N, 58°E	MR*	7.1988–6.1989	Ovezgeldyev <i>et al.</i> (1991)
Kyoto	35°N, 136°E	MR	5.1983–5.1984	Manson <i>et al.</i> (1985)
Atlanta	34°N, 84°W	MR	1974/75, 1976/77	Manson <i>et al.</i> (1985)
Kauai	22°N, 160°W	MF	10.1990–8.1992	Fritts and Isler (1994)
Punta Borinquen**	18°N, 67°W	MR	1977–1978	Manson <i>et al.</i> (1985)
Waltair	18°N, 83°E	MR*	7–8.1979	Devara <i>et al.</i> (1981)
Jamaica	18°N, 77°W	MR	3.1971–2.1972	Scholefield and Alleyne (1975)
Ramey	18°N, 67°W	MR	2.1981–6.1981	Roper (1984)
Christmas Island	2°N, 158°W	MR MF	1988–1989 1.1990–6.1991	Avery <i>et al.</i> (1989) Vincent (1993)
Mogadisho	2°N, 45°E	MR*	1968–1970	Babadjanov <i>et al.</i> (1974)
Jakarta	6°S, 107°E	MR	11.1992–10.1995	Tsuda (1995)
Townsville	20°S, 147°E	MF	1978–1980	Manson <i>et al.</i> (1985)
Grahamstone	33.3°S, 30°E	MR	1987–1993	Malinga and Poole (1997)
Adelaide	35°S, 138°E	MF	1978–1983 1984–1986	Manson <i>et al.</i> (1985) Manson <i>et al.</i> (1991)
Christchurch	44°S, 173°E	MF	6.1978–2.1980	Manson <i>et al.</i> (1985)
Mawson	68°S, 63°E	MF	1984–1986	Manson <i>et al.</i> (1991)
Molodezhnaya	68°S, 45°E	MR*	1967–1985	Portnyagin (1986)
Scott Base	78°S, 167°E	MF	12.1982–11.1984	Portnyagin <i>et al.</i> (1993)

Notations: \* - stations without height resolution; \*\* - only zonal wind component.

## MEAN ZONAL WIND

JANUARY

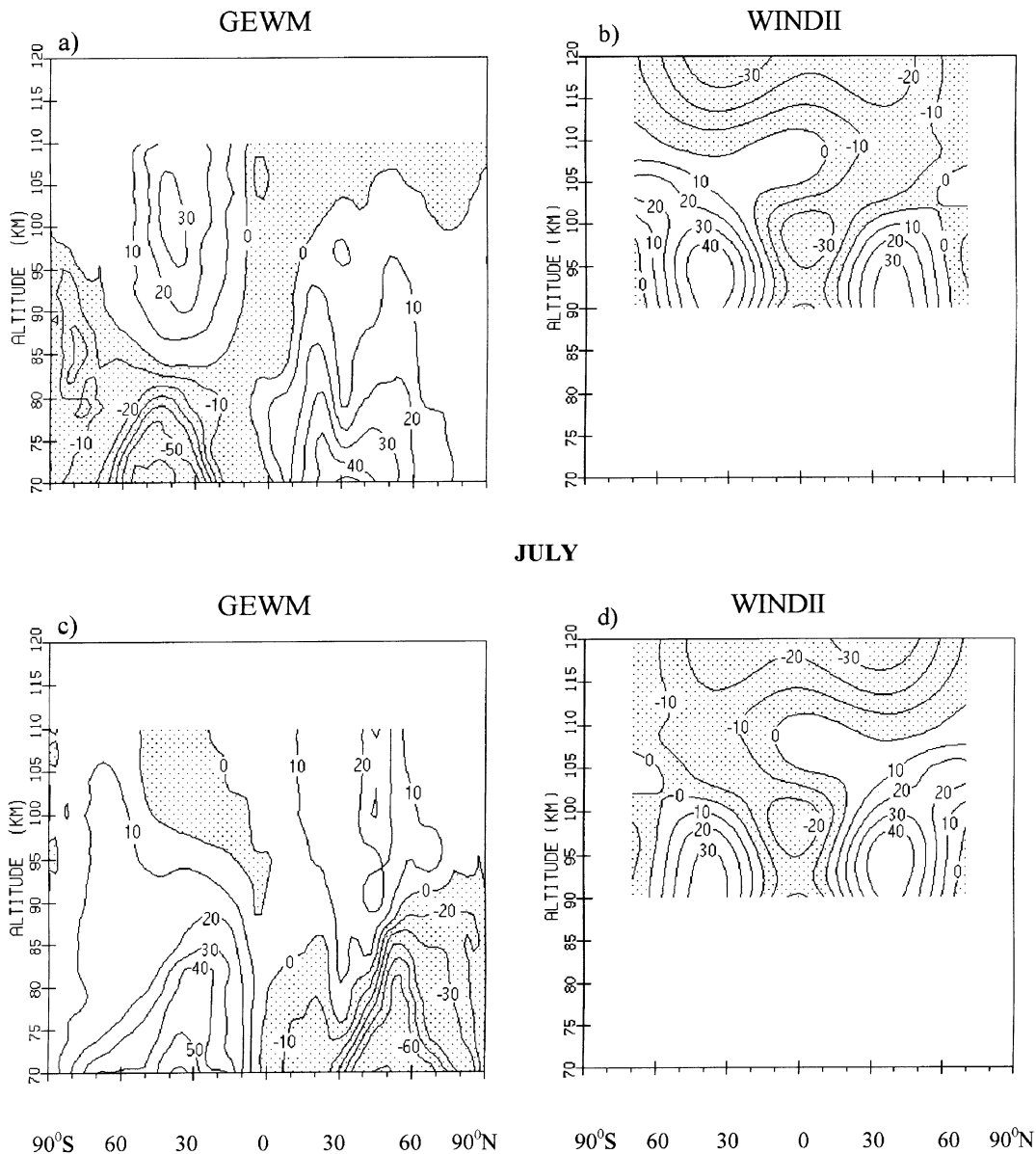


Fig. 1. Mean zonal wind for January and July. GEWM - Global Empirical Ground-Based Wind Model, WINDII - Empirical Model by Wang *et al.* (1997). Contour interval is 10 m/s. Westward winds are shaded.

The resulting monthly mean zonal wind values were calculated from the derived model for desirable height and latitude grids (usually,  $2.5^\circ$  in latitude and 1 km in height), and the height-latitude wind isoline plots were drawn with help of a proper software.

### 3. WINDII Model

The WINDII monthly mean zonal winds used in this analysis are retrieved from the empirical horizontal wind model presented by Wang *et al.* (1997). The model is constructed by using the total of 324 days of wind data inferred from green line airglow emissions observed by WINDII between 90 and

120 km during 1992 and 1993. The latitudinal coverage of the data alternates from  $72^\circ\text{S}/40^\circ\text{N}$  to  $42^\circ\text{S}/72^\circ\text{N}$  throughout the course of the year, determined by the WINDII viewing configuration relative to the satellite platform and the pointing direction of UARS which rotates  $180^\circ$  with a yaw period  $T_0 = 36$  days. The vertical resolution of the data is 3 km. Over the course of a day, the satellite sampling provides about 15 longitudinal points and two different local times at a given latitude. The two local times sampled daily at each latitude remain approximately constant, changing in fact by only 20 minutes per day for successive orbits. Two successive months of the data sets can achieve a full 24 hours of

## MEAN ZONAL WIND

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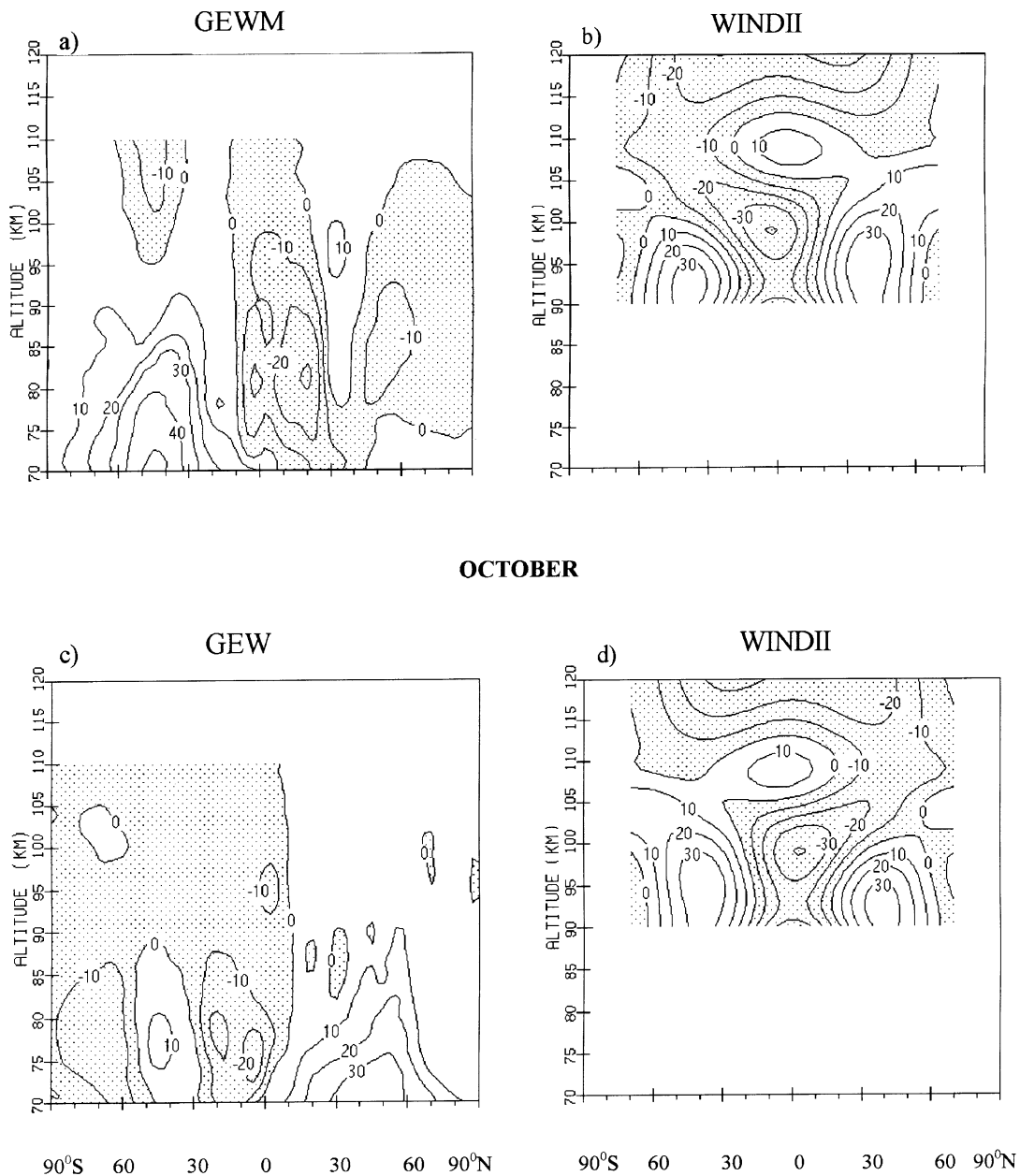


Fig. 2. Same as Fig. 1, but for April and October.

local time sampling. In accordance with the data characteristics, the vector spherical harmonic-Fourier expansion has been used to decompose the observed winds into mean winds, solar diurnal and semidiurnal tides, and stationary planetary wave components, as well as annual and semiannual variations in these fields. The model formulation is similar to that of HWM-90 and HWM-93 of Hedin *et al.* (1991, 1996), but greatly modified and extended to use the higher resolution global wind measurements now available from WINDII. With the spectral coefficients derived from the two years of WINDII wind data, the individual components of horizontal

winds for an arbitrary geographic location and time at the preset height levels can be reconstructed. This decomposition greatly reduces influences of tides and stationary waves on the model-retrieved monthly mean winds. More discussion and comparison between the model and the observed data can be found in the original paper.

#### 4. Monthly Mean Zonal Winds of GEWM and WINDII Models

In this section, we compare monthly mean zonal winds obtained from the GEWM and WINDII models. The results

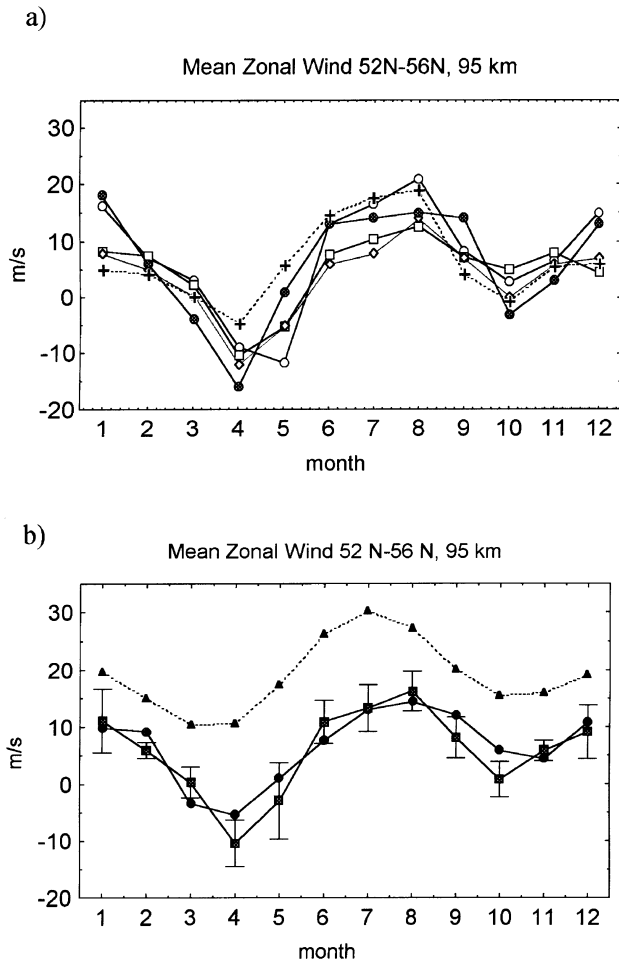


Fig. 3. Seasonal variations of mean zonal wind at 95 km. a) ○ - Kazan 56°N, □ - Obninsk 55°N, ◇ - Kùlungsborn 54°N, ● - Jodrell Bank 53°N, + - Saskatoon 52°N; b) ● - GEWM 55°N, ■ - experimental data, ▲ - WINDII 54°N.

on the meridional component will be published in a separate paper.

The height-latitude cross sections of the monthly mean zonal winds are shown in Figs. 1 and 2 for the solstitial and equinoctial conditions, respectively. The altitude coverages of the two models are different. Comparison of the model results is confined to the overlapping region of 90–100 km.

In order to examine the degree of consistency and difference between the two models, Figs. 3–5 show seasonal course of the zonal winds for some latitudes at the most reliable height of about 95 km, together with the complementary ground-based experimental data. From these figures, the seasonal courses of the GEWM and WINDII winds are seen to be very similar, but a systematic annual mean bias between the models is also evident. These observations are confirmed by Table 2, where the results of harmonic analysis for the zonal wind seasonal courses are presented. It is seen that the amplitudes and phases of the annual and semi-annual variations for two models are in good correlation. However, the significant offsets between the annual mean winds at different latitudes are also obvious.

The height-latitude structure of the offsets is shown in

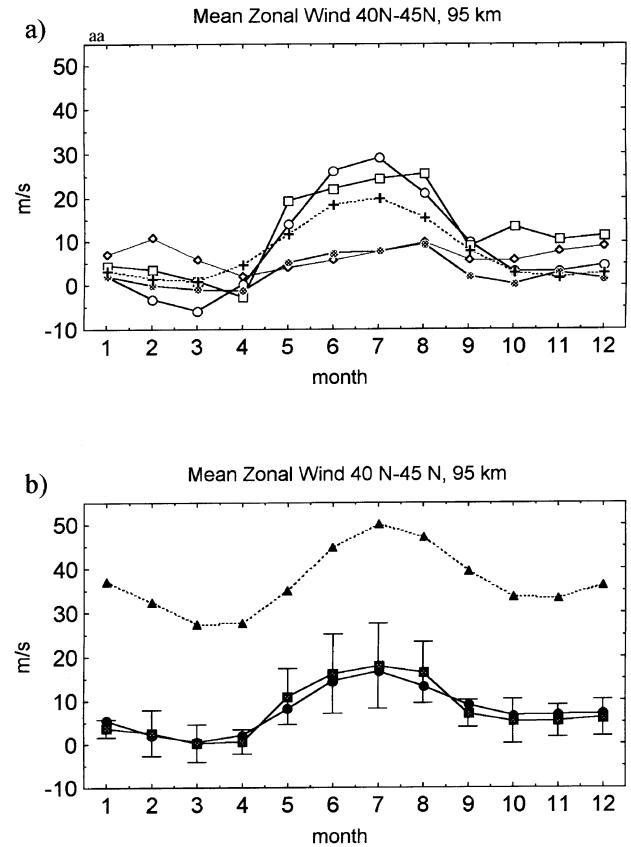


Fig. 4. Seasonal variations of mean zonal wind at 95 km. a) ○ - Monpazier 45°N, □ - Durham 43°N, ◇ - Frunze 43°N, ● - Jambol 42.5°N, + - Urbana 40°N; b) ● - GEWM 42.5°N, ■ - experimental data, ▲ - WINDII 42.5°N.

Fig. 6a. The systematic bias between GEWM and WINDII model is seen to have a very regular latitudinal structure, which practically does not depend on height. As a result, a very good agreement between the models could be obtained, provided that the WINDII annual wind values were subtracted by a term of  $A \cos 4x$ , where  $x$  is colatitude and  $A$  is about 20–25 m/s and is practically independent of height. This simple function is determined from the offsets at a given height level by the least-squared fitting.

It is interesting to examine the height-latitude structure of annual mean wind itself. The annual mean zonal winds derived from GEWM (Fig. 6b) exhibit little height dependence, but show a strong latitude dependence of  $A' \cos 4x$ , where amplitude  $A'$  may be roughly considered as independent of height. The certain similarity between the Fig. 6a and Fig. 6b is obvious. Then, our analysis shows the annual mean winds of WINDII can be expressed as  $A'' \cos 4x$ , where height-independent amplitude  $A'' = A + A'$ , and is in average equal to  $(2-2.5)A'$ . Thus, the offset between the annual mean winds for the two models (cf. Figs. 6a and 6b) can be attributed to a difference by about 2–2.5 times in the relative wind magnitude.

Due to nearly the same annual and semi-annual variations and large annual mean bias for the two models, the monthly mean winds in Figs. 1 and 2 exhibit some similarities and significant differences. For the winds in January (Fig. 1 top),

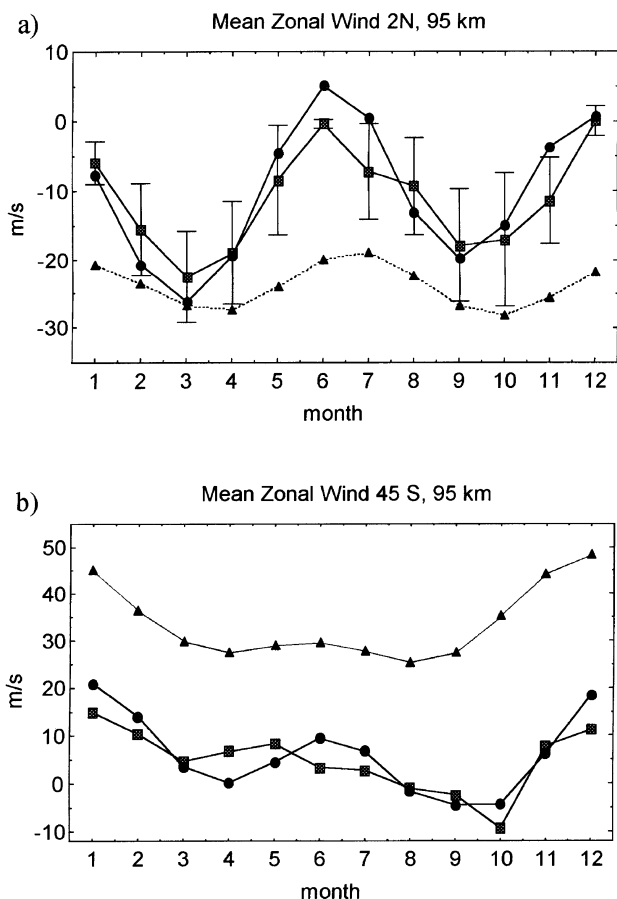


Fig. 5. Seasonal variations of mean zonal wind at 95 km. a) ● - GEWM 2.5°N, ■ - experimental data (2°N), ▲ - WINDII 2.5°N; b) ● - GEWM 45°S, ■ - experimental data (44°S), ▲ - WINDII 45°S.

the most important features of the height-latitude variations are generally common for the two models.

At the mesospheric heights, the dominating structures are caused by the well-known strato-mesosphere winter cyclon and summer anticyclon vortices. In the lower thermosphere, the wind reversal with height is observed at mid-latitudes, but the height of this reversal is not the same for the two models. At lower latitudes, westward winds are observed in the whole height range. However, the exact position and the maximum wind amplitudes of these main structures are different.

The circulation patterns in July (Fig. 1, bottom) are practically the same as in January, but are reversed relative to the equator in comparison with January according to changing of seasons (winter-summer). Again the localization of the dynamical structures and the wind amplitudes are different for the different models.

For April and October (Fig. 2), the degree of similarity between the models is not too high as that for solstice months. The equinoctial GEWM winds show much weaker amplitudes in comparison with those of the space-based model.

## 5. Discussion

We have found significant systematic difference in annual mean zonal winds between the GEWM and WINDII empir-

Table 2. Zonal prevailing wind seasonal course components 95 km.

	A0	A1	Ph1	A2	Ph2
52°N–56°N					
GEWM 55°N	6.6	5.6	8.9	6.2	.7
Mean experiment	5.7	5.4	8.8	8.3	.6
WINDII 54°N	18.9	6.0	7.3	6.0	.4
40°N–45°N					
GEWM 42.5°N	7.7	5.8	7.2	3.7	.1
Mean experiment	7.7	7.0	6.8	4.5	.1
WINDII 42.5°N	37.0	7.2	7.3	6.5	.4
2°N					
GEWM 2.5°N	−10.2	4.1	7.4	13.3	5.5
Mean experiment	−11.2	1.7	7.9	9.4	5.8
WINDII 2.5°N	−23.8	1.0	5.7	4.1	0.2
44°S					
GEWM 44°S	6.2	7.1	1.1	8.7	0.1
Mean experiment	5.0	7.2	0.9	8.7	0.2
WINDII 45°S	33.8	9.4	11.7	5.2	5.5

A0 - annual mean wind (m/s), A1 - amplitude of annual harmonic (m/s), Ph1 - phase of annual harmonic (month of maximum), A2 - amplitude of semiannual harmonic (m/s), Ph2 - phase of semiannual harmonic (month of maximum).

ical models, while the annual and semi-annual variations of zonal wind field in the two models are in good agreement, respectively. In view of the general consistency between the model representations and their own observational datasets based on which the models are constructed, the revealed discrepancy is believed to reflect the deviation between the wind measurements by the ground and space techniques.

We note that there is a lot of direct intercomparison between ground-based and space-based data. Intercalibration of HRDI and WINDII wind measurements by Burrage *et al.* (1997) have shown that there is no significant wind-speed bias between the two datasets. The directions are also in good agreement; the WINDII winds appear to have an offset in the zonal component of about  $-6$  m/s relative to the HRDI winds, which in turn have an offset of  $+2$  m/s relative to MF radars and rockets measurements, suggesting an absolute zonal wind bias of  $-4$  m/s for WINDII.

The results of the most comprehensive intercomparison between simultaneous MLT winds measured by rockets, ground-based MF and meteor radars, and by space-based HRDI/WINDII were reported by Burrage *et al.* (1996) and Hasebe *et al.* (1998). Burrage *et al.* (1996) show that the HRDI winds agree well to those observed by radars, rockets and WINDII, except that the relative magnitude of wind speed is smaller in MF radars than in HRDI. Hasebe *et al.* (1997) concluded that the accuracy of HRDI winds relative to meteor radars, situated in Shigaraki and Jakarta, is found to be high, in that overall differences in the wind Shigaraki and Jakarta, is found to be high, in that overall differences in

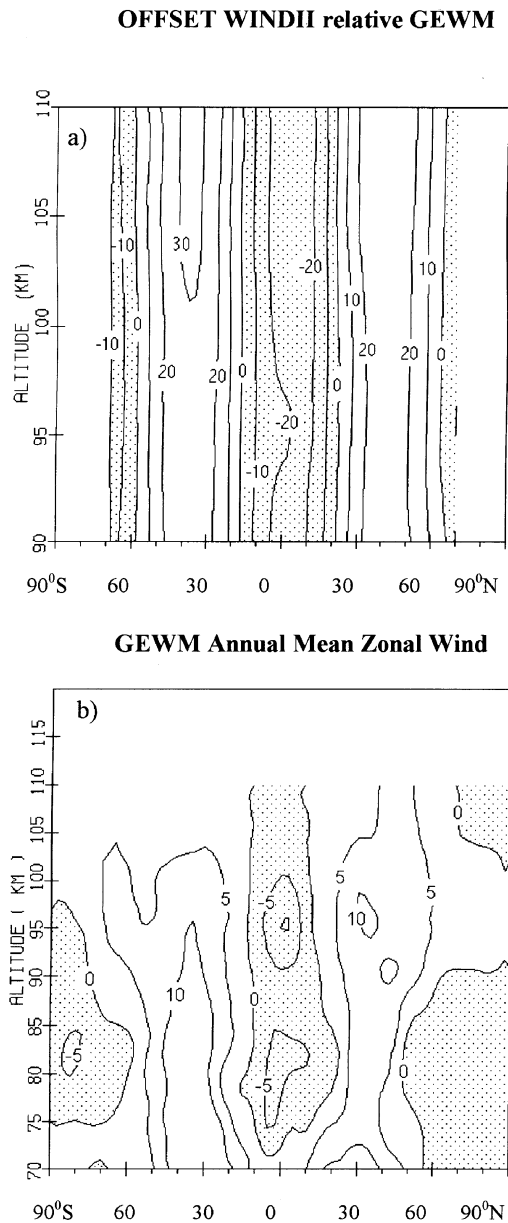


Fig. 6. a) High-latitude structure of the systematic difference between WINDII and GEWM annual mean winds. Contour interval is 10 m/s. b) High-latitude structure of GEWM Annual Mean Wind (m/s). Contour interval is 5 m/s. Westward winds are shaded.

the wind velocities are not statistically significant at the 0.01 level between HRDI and meteor radars.

Most of the above mentioned comparisons, however, has been made for instantaneous wind profiles. In spite of the conclusions drawn from the comparison between HRDI and Jakarta meteor radar data, we note that there is a constant strong shift in the monthly mean values, while the individual wind profiles show very good correlation. This is of particular interest to our comparisons for monthly and annually mean winds.

Actually, the GEWM is mainly based on the meteor radar wind measurements, especially at about 95 km. As shown in Section 4, significant systematic differences between the GEWM and WINDII are definitely exist. Long-term meteor

and radar measurements (see Table 1) at different geographical sites have shown that the main regularities of the lower thermosphere wind regime are very persistent. For example, at 95 km level and at the high and mid-latitudes above 45 degrees in the Northern hemisphere, all meteor and MF radar wind data since the very old meteor data (1957–1961) from Jodrell Bank and up to now showed the same seasonal behavior: eastward flow for winter and summer seasons with a rather strong wind reversal in spring time and with a weaker wind, decreasing in fall. For more than 40 years of observations at different longitudes we have not found any exception from this general rule. However, the space-based WINDII empirical model and HRDI winds do not show any spring zonal wind reversal in this latitudinal belt, due to the large offset in annual mean winds.

Let us consider possible origin of the observed deviations between ground-based and satellite-based MLT winds. The ground-based radar measurements were independently taken at the different sites with the different techniques over decades of observation period. Inter-annual and longitudinal variability in the ground-based wind data may play a certain role in the observed differences. We have estimated the mean square deviations stipulated by the inter-annual and longitudinal variability. The values do not exceed 5–7 m/s at the high and mid-latitudes (cf. Figs. 3, 4) and are about 10 m/s at the lower latitudes (cf. Fig. 5). This deviation cannot reconcile the observed discrepancy between the annual means of the GEWM and WINDII winds. More importantly, as shown in Fig. 6a, the systematic bias remarkably regularly varies with latitude, but practically does not depend on altitude. This height-independent feature cannot be attributed to the inter-annual and longitudinal variabilities of the ground-measured monthly means, since such variations are expected to have height distribution. In fact, if the latitudinally varying bias were eliminated in a way suggested in Section 4, a high consistency between the GEWM and WINDII winds could be achieved. This provides an indirect confirmation that the inter-annual and longitudinal effects on the bias are unimportant.

There are several factors which may be responsible for the detected bias. Vincent *et al.* (1994) have shown that in cases of severe MF receiver saturation the observed wind velocities would be systematically smaller than the actual wind velocities. Meek *et al.* (1997) have pointed out that some speed-ratio bias in the sense  $MFR < HRDI$  is expected on statistical grounds. These discussions may explain the difference between MF radar and HRDI/WINDII data. However, most of GEWM data are taken by meteor radars, which are believed to be reliable at least at the most significant height of about 95 km.

Palo *et al.* (1997) discussed a possibility that a possible reason for the bias between ground-based and satellite-based data may be connected with a long-period intraseasonal oscillations which would be suppressed in the UARS climatology. Other possible origins may be attributed to the systematic errors in the absolute wind measurements by the satellite-based optical techniques. The WINDII winds are derived from airglow observations. In the inversion calculations, the airglow intensity error of 5% may result in the wind error of 20 m/s. This is likely a candidate (Y. J. Rochon, private

communication, 1998), though the exact nature is to be investigated.

For HRDI, one possible origin is the systematic errors in the absolute calibration of the line shift to Doppler velocity factor on board of UARS. As pointed out by Burrage *et al.* (1996), if this factor were substantially in error, the correction to account for the Earth rotation effect would lead to large errors in the measured winds, due to the maximum Earth rotation velocity of 463 m/s at the equator in zonal direction, larger than typical MLT winds by an order of magnitude. Since the correction depends on viewing direction and latitude, the wind errors would vary with latitude, but be independent of height, suggesting the same behavior as that observed in our analysis. An additional evidence can be found in figure 24 of the above-referred paper, where an example shows the relationship between the Earth rotation contribution and the total velocity HRDI measurements for the different azimuths. The average slope of the relationship is very close to 1 (for many days of HRDI data, this slope is nearly always in the range 0.95 to 1.05). As pointed out by the authors, if the line shift to wind factor was in error, then this slope would not be 1. Actually, this slope significantly deviates from the average  $y = x$  line in the different latitudinal belts. The deviations have a wavelike structure relative to the latitude, similar to those of the above-discussed bias. Therefore an uncertainty in the correction to account for the rotation of the Earth may cause the observed differences between ground-based and space-based MLT wind models.

## 6. Conclusions

The GEWM and WINDII empirical wind models are constructed from the meteor/MF radars and WINDII green line airglow measurements, respectively. Comparison between the zonal winds retrieved from the models has revealed a general consistency, in particular, almost the same annual and semi-annual variation components in global scale wind structures.

However, systematic bias exists in the annual mean zonal winds. The significant offsets result in eastward flow throughout the year at 95 km at the high and mid-latitudes in the Northern hemisphere, inconsistent with the well-known from ground-based measurements zonal wind reversal in that region during equinoxes. The discrepancy is unlikely associated with model representation and/or inter-annual/longitudinal variabilities of the observational datasets, since similar deviations are also observed in HRDI data. Possible origin responsible for the detected bias are suggested, more investigations are required.

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