Rock magnetic study of fluvial Holocene soil from Buenos Aires province (Argentina)

Carlos A. Vasquez^{1,2} and Hugo G. Nami^{2,3}

¹Ciclo Básico Común, Universidad de Buenos Aires. Ciudad Universitaria (Pabellón III), (1428) Buenos Aires, Argentina

²CONICET—Instituto de Geofísica "Daniel A. Valencio", Departamento de Ciencias Geológicas, Facultad de Ciencias Exactas,
Físicas y Naturales, Universidad de Buenos Aires, Ciudad Universitaria (Pabellón II), (1428) Buenos Aires, Argentina

³Associate Researcher, Deparment of Anthropology, National Museum of Natural History,
Smithsonian Institution, Washington D.C., 20560, U.S.A.

(Received December 22, 2005; Revised June 4, 2006; Accepted June 5, 2006; Online published November 8, 2006)

The magnetic characteristics of soils are widely used in environmental and paleoclimatic investigations for studying the several factors involved in the soil formation process. We propose here a new analytical tool that takes into account the variations in magnetic properties correlated with grain sizes and concentrations of ferrimagnetic minerals. This analytical tool is based on a mathematical model of well-established magnetic properties in samples of known grain sizes and was used in this study to determine changes in the grain size and concentration of ferrimagnetic minerals along a terminal Pleistocene/Holocene fluvial section located in the northeast of Buenos Aires province. These variations may reflect a humid period prevailing in the area and may be associated with climate changes that occurred in the Chaco-Pampean region during the Middle Holocene.

Key words: Rock magnetism, Magnetite, Titanomagnetite, Holocene, Buenos Aires.

1. Introduction

The magnetic characteristics of soils are widely used as defining parameters in environmental and paleoclimatic studies (Orgeira et al., 1998, 2003; Jordanova and Jordanova, 1999; Petrovsky et al., 2001; Thomson et al., 2001). Several climatic and environmental factors have been identified as contributing to the magnetic signal of soils, including temperature, humidity, winds, floods, rivers, and ground water table, among others (Verosub and Roberts, 1995; Maher and Thompson, 1999). More specifically, temperature and humidity can be associated with the pedogenetic process involved in the production of SP (superparamagnetic) or SP/SSD (stable single-domain) boundary particles (Dearing et al., 1996; Jordanova and Jordanova, 1999), while winds and rivers are related to the transport process of detritic magnetic particles. In this case, grain size ranges between those of PSD (pseudo single-domain) and MD (multi-domain) boundary particles (Evans, 2001; Matasova et al., 2001). The influence of the ground water table is more complex because magnetite is formed in this interface (MacBride, 1994), and changes in pH and redox can dissolve the magnetite grains (Jordanova and Jordanova, 1999).

In soils formed from wind parental material, such as loess-paleosoil sequences, the magnetic signal results from two opposite formation processes—wind transport and pedogenesis—with the balance between the two being determined by the distance from the parental detritic material

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source (Hunt *et al.*, 1995; Petrovsky *et al.*, 2001; Evans, 2001). Although the (changing) transport mechanism of a river is known to affect the magnetic signal of fluvial material, it is also possible the basis of this association lies in the fluvial parental material (Jordanova *et al.*, 1997).

The aim of this investigation reported here was to estimate the grain sizes and concentrations of ferrimagnetic minerals in a fluvial section located in Buenos Aires province (Argentina), based on a mathematical model of well-established magnetic properties of samples of known grain size (Dunlop and Özdemir, 1997).

2. Sampling Site and Chronological Considerations

The Lomas del Mirador (LM; 34°39.29′S, 58°32.17′W) site is located in the outskirts of Buenos Aires City (Fig. 1). From a geomorphological and geological viewpoint, Buenos Aires belongs to the Chaco-Pampean region (Russo et al., 1979). There is to date no evidence of fluvial activity at LM, however, the sediments belong to a characteristic floodplain deposit. This area was an active fluvial environment during the terminal Pleistocene/Holocene and also subject to pedogenesis. There is a currently fully developed soil that is suggestive of a period of non-deposition and landscape stability (Holliday 1985; Kraus and Brown 1986). According to Favier Duvois (personal communication), the deposit at LM is similar to the Guerrero member of the Luján formation, which is a useful horizon marker for late Pleistocene and early Holocene sediments (Cione and Tonni 1995; Tonni et al., 1999, 2003).

A $1-m^2 \times 2.7$ -m deep trial pit in the sediment deposit revealed the presence of five stratigraphic levels (I–V; Fig. 2):

	Material	Depth	Method	Absolute date	Relative minimum	Identification
	dated			(years BP)	date (years BP)	number
_	Mollusc	0.60/0.70	¹⁴ C - AMS	Modern	_	CURL-5502
	Mollusc	1.20-1.25	¹⁴ C - AMS	4900 ± 110	_	OS-24330
	Sediment	0. 55	OCR	_	115	Act # 2943
	Sediment	1.22	OCR	_	789	Act # 2944
	Sediment	1.55	OCR	_	1374	Act # 2945
	Sediment	1.88	OCR	_	3110	Act # 2946
	Sediment	2.05	OCR	_	4546	Act # 2947
	Sediment	2.30	OCR	_	5525	Act # 3230
	Sediment	2.30	OCR	_	6065	Act # 3360
	Sediment	2.39	OCR	_	6139	Act # 3228
	Sediment	2.39	OCR	_	6609	Act # 3362
	Sediment	2.39	OCR	_	6836	Act # 3234
	Sediment	2.63	OCR	_	7165	Act # 3363
	Sediment	2.63	OCR	_	7556	Act # 3229

Table 1. List of dates obtained from LM deposit. Depths are given in meters below the surface.

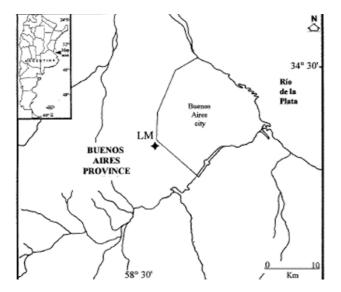


Fig. 1. Location map of the sampling site (LM) in Northeastern Buenos Aires province.

I is an artificial landfill, II is very dark brown (2/2, 10 YR Munsel color chart) highly compacted silt sediments, III is a dark-brown (2.5/3, 7.5 YR) silty clay level, IV is a palebrown (6/3, 10 YR) silty sand sediment, and V is a light yellowish brown (6/4, 10 YR) silty clay deposit with calcareous inclusions. Levels II-IV belong to the same soil formation horizon (Nami, 2006). After the northern profile of the trial pit had been cleaned, approximately 100-g samples (n = 19) were collected in nylon bags. Directional analysis was carried out on those levels belonging to the terminal Pleistocene and Holocene (Nami, 2006). Hence, consecutive samples were continuously taken without intervals in levels II–IV between 0.54 and 1.49 m below the surface.

The age of the deposit was determined using different dating methods. Two ¹⁴C dating analyses on shell samples using mass accelerator spectrometry (AMS) revealed that a small piece of unidentified mollusk from level II contains more ¹⁴C than the 1950 atmosphere; this sample is therefore "modern"—<1950 AD—which is coincident with the age of the artificial landfill. ¹⁴C dating of a specimen of the

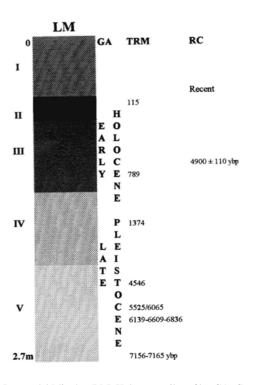


Fig. 2. Lomas del Mirador (LM) Holocene soil profile. GA: Geological Age. TRM: Mean residence time. RC: Radiocarbon age.

micro-faunal mollusk (*Mactra isabelleana*) belonging to the lower part of level III revealed an age of 4900±110 years BP (OS-24330), which is coincident with radiocarbon dates obtained from M. isabelleana collected from other sites in the region (Tonni *et al.*, 1999). This micromollusc is a typical fossil found in Holocene deposits from Buenos Aires (Aguirre, 1988, 1993).

The mean residence time (MRT), of the soil, which is a measure of the mixing of the young organic carbon with organic carbon from earlier stages of pedogenesis (Stein, 1992), was determined employing the oxidizable carbon ratio (OCR) analytical method (Frink, 1992, 1994, 1995; Scharpenseel, 1971). Twelve OCR dates were obtained from levels II–V ranging between 115 and 7556 years BP (0.11–7.55 kybp; Table 1). Taking into account the radiometric date, these dates must be considered as relative mini-

mum ages. Although there are notably differences between the ¹⁴C and OCR dates that are correlated with depth, the latter dates simply confirm that the soil surface was open to organic input for a long period (e.g., Scharpenseel, 1971; Stein, 1992). Our MRT values therefore suggest that the deposit was open to organic material deposition during the last ~8 kyr (i.e. through much of the Holocene), thereby confirming the terminal Pleistocene/Holocene age of the soil formation. In fact, its boundary with the Pleistocene was conventionally established by Dawson (1992) to be 10 kyr BP. In addition, a comparison of the results obtained in other sections from Argentina with the paleomagnetic directions registered at LM suggest that the magnetic signal of the latter belong to the terminal Pleistocene-Holocene (Nami, 2006). To summarize, based on geological, chronological, and paleomagnetic studies it is possible to suggest that the magnetic signal reported in this paper belongs to the Holocene.

3. The Mathematical Model of Grain Size and Concentration

Magnetic properties of a soil are correlated with sizes and concentration of the grains in a qualitative manner. While it is difficult to establish a mathematical relationship between grain size and magnetic properties for most minerals (Dunlop and Özdemir, 1997), magnetite PSD grains can be easily modeled. We therefore propose a mathematical model based on well-established data that is applicable to the quantification of grain sizes and their concentrations in natural samples where PSD magnetite drives the magnetic signal.

3.1 Relationship between M_{rs}/M_s and the grain size

The ratio M_{rs}/M_s (specific saturation remanence/specific saturation magnetization) is the higher sensitivity parameter to magnetic grain size of magnetite in the PSD-MD range (Dunlop and Özdemir, 1997; Dunlop, 2002). Both M_{rs} and M_s are grain size- and concentration-dependent parameters, but their ratio is a concentration-independent parameter (Banerjee, 1981; Benerjee and Hunt, 1993). The theoretical relationship between the grain diameter d and the M_{rs}/M_s ratio, as proposed by Halgedahl and Fuller (1983) is:

$$M_{rs}/M_s = A * \exp(-(d/d_o)^{1/2})$$
 (1)

where A and d_o are constants that are dependent on magnetite or titanomagnetite characteristics (Hartstra, 1982; O'Donovan *et al.*, 1986; Dunlop and Özdemir, 1997; Dunlop, 2002). From this, it is possible to determine the grain size if the M_{rs}/M_s ratio is known.

A comparison of Eq. (1) to the experimental data of Hartstra (1982) and O'Donovan *et al.* (1986) is shown in Fig. 3. This figure reveals the presence of a very good agreement between the data of Hartstra (1982) and the theoretical model, while there is more dispersion between the data of O'Donovan *et al.* (1986) and the theoretical model, although the fit is still reasonable good. The parameters for the model are:

$$M_{rs}/M_s = A_o + A_1 * \exp(-(d^{1/2} - x_o)/t_1)$$
 (2)

where $A_o = 0.0512$, $A_1 = .46148$, $x_o = 1.41421 \mu m^{1/2}$, $t_1 = 1.89716 \ \mu m^{-1/2}$.

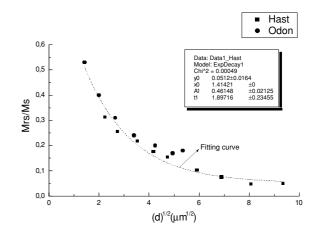


Fig. 3. Fiting of Hastra (1982) and O'Donovan *et al.* (1986) data of titanomangetites to an exponencial function of M_{rs}/M_s vs. square root of diameter.

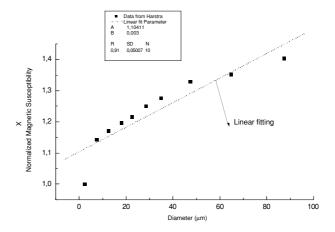


Fig. 4. Linear fit of diameter vs. susceptibility, data from Hastra (198).

Arranging terms in Eq. (2) results in the following:

$$M_{rs}/M_s = A_o + A_1' * \exp(-(d/t_1^2)^{1/2}),$$
 (3)

where $t_1^2 = d_o$, and $A'_1 = A_1 * \exp(x_o/t_1)$; A_o is an asymptotic term, indicating that the M_{rs}/M_s tends towards a constant value when d goes to infinity.

From Eq. (2), it is possible to determine the titamomagnetite grain diameter d, according to the expression:

$$d = (x_o - t_1 * \ln((M_{rs}/M_s - A_o)/A_1))^2$$
 (4)

The model of Halgedahl and Fuller (1983) can be used in sediments where titanomagnetite or magnetite is the main magnetic carrier and the grain diameter ranges between 5 and $100 \mu m$. Both of these conditions remain valid for the soil under study here where magnetic detritic grains can be transported by rivers or wind (Scasso and Limarino, 1997).

3.2 Magnetic susceptibility and titanomagnetite concentration

Changes in the ferrimagnetic initial magnetic mass susceptibility χ can be attributed to two main causes: the concentration and grain size of magnetic minerals. At the same grain size, susceptibility χ is proportional to concentration (Dunlop and Özdemir, 1997). According to Hartstra (1982) and O'Donovan *et al.* (1986), there is a relationship between

Depth (cm)	K (m ³ /kg) LF	K (m ³ /kg) HF	B_{cr} (mT)	B_c (mT)	M_{rs} /masa (1E-8)	M _s /masa (1E-8)
-54	1.262E-06	1.214E-06	29.6	10.9	16.546875	102.34375
-58	1.2785E-06	1.225E-06	29.6	10.2	15.81481481	97.96296296
-63	1.452E-06	1.386E-06	30	10.5	17.75384615	108.9230769
-68	1.6069E-06	1.532E-06	30.2	10.58	18.22807018	113.8596491
-73	1.8476E-06	1.769E-06	32.4	10.5	22.68292683	148.7804878
-78	1.9877E-06	1.903E-06	32.7	10.99	27.3442623	180.442623
-83	1.9804E-06	1.903E-06	33.1	10.96	27.33962264	182.0754717
-88	2.0201E-06	1.932E-06	32.9	11.1	27.35185185	183.444444
-93	2.0641E-06	1.978E-06	34.1	11.7	28	180.3833333
-98	2.0399E-06	1.955E-06	33.5	11.5	27.44827586	177.5862069
-103	1.9789E-06	1.903E-06	34.5	11.79	30.03333333	196.9
-108	1.9864E-06	1.913E-06	34.3	11.57	27.3968254	181.3174603
-113	2.0398E-06	1.967E-06	33.7	11.5	31.43396226	203.5849057
-118	2.0621E-06	1.986E-06	34.4	11.81	33.16981132	215.7169811
-123	1.9817E-06	1.906E-06	34.7	12.2	28.57377049	186.7213115
-128	1.9532E-06	1.886E-06	34.3	11.37	27.15789474	185.4385965
-133	1.933E-06	1.885E-06	34.6	12	31.82978723	213.1914894
-138	1.9301E-06	1.89E-06	34.4	11.53	30.33333333	206.4509804
-143	1.8907E-06	1.845E-06	34.4	11.59	28.375	190.5178571
-148	1.9002E-06	1.859E-06	34.4	11.59	28.375	190.5178571

Table 2. Mass Susceptibility at 470 Hz (K LF), and 4700 Hz (K HF), Coercitive Magnetic Field (B_c), and Coercitive Remanence Magnetic Field (B_{cr}), Remanent Saturation (Mr) and Saturation of Magnetization (M_{rs}) in LM profile.

grain size and the susceptibility of magnetites and titanomagnetites in the range 5–100 μ m. Consequently, it is possible to determine a linear correlation between these based on their data, as shown in Fig. 4.

In order to analyze the effect of grain size on susceptibility, a normalized susceptibility is defined:

$$\chi_N(d) = \chi/\chi_L \tag{5}$$

where χ_L is the lowest susceptibility value. The relation between normalized susceptibility $\chi_N(d)$, and grain diameter d is then:

$$\chi_N(d) = A + B * d \tag{6}$$

where the A and B parameters are shown in Fig. 4.

With this susceptibility "corrected" by grain size, it is now possible to determine the magnetic concentration if the magnetic mineralogy is known by independent tests: coercitive magnetic fields (H_c and H_{cr}), χ vs. temperature curves, and high and low temperatures, among others.

Let χ_N be the normalized susceptibility to the lowest susceptibility by mass unit; N(d), the grain concentration, which is diameter (d)-dependent with a range between 0 and 1; χ_d , the bulk susceptibility from grains of diameter d per mass unit. If all of the grains of the sample have the same diameter and all are magnetic grains, N=1, and the normalized bulk susceptibility χ_N is

$$\chi_N = \int N(d)\dot{\chi}_N(d)dd = A. \int N(d)dd$$

$$+ B. \int N(d).ddd \approx A.N$$
(7)

because the *B* factor is threefold lower that *A* and $\int N(d).ddd$ is acoted ($0 \le N(d) \le 1$, $d \le 100$ E-6 m). The bulk normalized susceptibility χ_N is mainly driven by the concentration factor *N*. It is only a rough model that may be used in a semi-quantitative analysis.

4. Magnetic Analysis and Grain Concentration Determination

4.1 Experimental methods

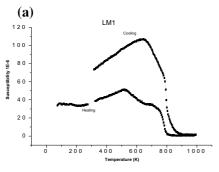
In the laboratory, sub-samples from the bagged samples were air-dried at room temperature, crushed in agar mortar and mixing several times in order to have representative samples. They were then weighed and packed into plastic boxes (2.5×2.5 cm; height×diameter) for magnetic analysis. The remaining portions of the samples were used for non-magnetic treatments.

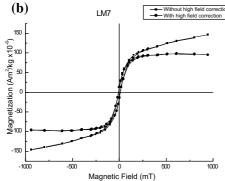
Magnetic susceptibility was measured in the Bartington MS2 analyzer. A Molspin VSM Nuovo vibrating sampling magnetometer was used to measure the M_s , M_{rs} , B_c and B_{cr} magnetic parameters at room temperature. Measurements at high temperatures were carried out using a Bartington MS2W susceptibilimeter.

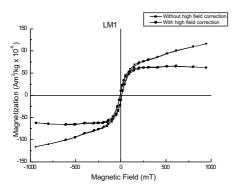
4.2 Results

Thermomagnetic analysis, hysteresis cycles, and IRM were carried out. The thermomagnetic curves are shown in Fig. 5. and the hysteresis and IRM parameters are shown in Table 2. The thermomagnetic curves show that the susceptibility is near constant, in the low temperature zone, without any Verwey transition. On the other hand, the high temperature zone shows a peak near 500 K, after which the curve is very similar to that of magnetite, with a marked drop near the magnetite Curie point (approx. 850 K). The susceptibility curve does not show the characteristic Hopkinson peak of SSD magnetite near the Curie point (Dunlop and Özdemir, 1997). These two results together with the flat curve at the low temperature are in agreement with a partial oxidation of magnetite (Dunlop and Özdemir, 1997; Thompson *et al.*, 2001; Carter-Stiglitz, 2002).

The susceptibility and hysteresis parameters of 19 levels in the LM section are shown in Table 2. The χ percentage difference in susceptibility at low and high frequencies is lower than 2%, indicating that the SP fraction is very low. This result is in agreement with the detritic origin of magnetic minerals because under anoxic conditions a pedoge-







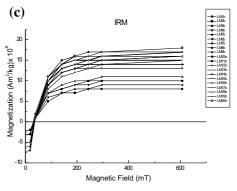


Fig. 5. (a) Heating and cooling susceptibility in air of LM1 sample. The Verwey transition is not present in the low temperature section. It is indicative of oxid magnetite as the main magnetic carrier. Mineralogical transformation of titano-magnetite to magnetite is shown as a drop in heating curve at 850 K, near the Curie point. (b) Hysteresis loops from two representatives samples, LM1 and LM7. (c) Isothermal Remanent Magnetization curves from all the samples. An initial magnetization at 800 mT is given and after a reverse magnetic field is applied.

netic neoformation of SP magnetite would occur (Petrovsky et al., 2001).

The median values of B_{cr} and B_{c} (±standard error) shown in Table 3 are in agreement with a homogeneous magnetic mineralogy of similar grain size. The results pre-

Table 3. Media value and Standard Error of Media for the magnetic parameters displayed in Table 1.

Magnetic Parameter	Media Value	Standard Error		
(units as in Table1)		of Media		
B_{cr}	33.09	0.40111		
$B_{\scriptscriptstyle C}$	11.294	0.1231		
M_{rs}	26.25946	1.16286		
M_s	172.30797	8.3062		
KLF	1.85977E-6	5.61095E-8		
KHF	1.79186E-6	5.46878E-8		

Table 4. Magnetic properties of several magnetic minerals. References: 1) Harstra (1982), 2) O'Donovan and colleagues (1986)

Substance	χ (m ³ /kg)	M_{rs}	M_{s}	B_{cr}	B_c	Ref.
	1E-6	(Am ² /kg)	(Am²/kg) (mT)	(mT))
Magnetite (sintetic)	0.824	0.028	0.16	25.7	12.3	2
Maghemite (sintetic)	0.544	0.029	0.106	27.3	15.4	2
Magnetite (natural)	0.58	0.018	0.121	32.1	12.8	2
Titanomaghemite (natural)	0.198	0.018	0.07	53.4	24.8	2
Magnetite 150–250 μm	601.9	0.9		25		1
Magnetite 25–30 μ m	686.12	2.31		33		1
Magnetite 5–10 μ m	529.04	11.07		54		1
Titanomagnetite 25–30 μ m	284	0.84		8.5		1
Titanomagnetite 5–10 $\mu \mathrm{m}$	196	1.74		11.5		1
Maghemite 25–30 μ m	361.91	3.74		31		1
Maghemite 5–10 μ m	317.93	3.8		37		1
Hematite 150–250 μ m	1.277	0.215		59		1
Hematite 25–30 μ m	1.576	0.234		72.5		1
Hematite 10–15 μm	1.178	0.231		140		1

sented in Table 4 indicate that that choice of natural magnetite grain sizes falls in the range 25–30 μ m.

The standard median error shown in Table 3 is nearly threefold higher for M_s , M_{rs} , and K than for B_{cr} and B_c . However, as B_{cr} shows a 13% change along the profile and B_c , a 9% change, these parameters can be related to minor increases in the grain sizes of the magnetite. The changes in M_{rs} is 42%, in M_s 46%, for both K 34%, indicating that variations in these magnetic grain concentration parameters are larger than those of magnetic grain size (B_c and B_{cr}) and that they control the magnetic signature in these sediments.

Figure 6(a) illustrates the semi-quantitative model application along the LM profile, depicting oscillations around 70 cm and between 90 cm and 103 cm. However, the general trend is an increase of grain size from the surface downwards. This is in agreement with the proposed transport model, where flooding events can be related to oscillations in grain size.

Based on the experimental data variations in the concentrations of magnetic minerals are influenced by grain size. The mathematical model applied to this soil (Fig. 6(a)) can be used in order to "clean" the magnetic record and to extract the signal based on concentration only. A simple inspection of the plot of the concentration parameters M_{rs} , M_s , and χ (Fig. 6(b)) does not reveal minor changes in concentration because they are also grain size-dependent. A comparison of Fig. 6(a) and (b) shows a peak near -90 cm (Fig. 6(a)) that can be attributed to a flooding event. On the other hand, the grain size is nearly constant along the

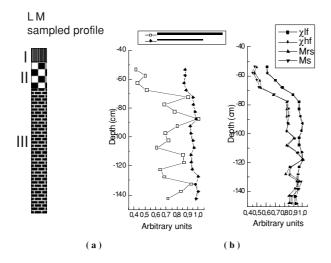


Fig. 6. (a) Normalized grain concentration and grain size changes with depth. Variations in grain size are lower than grain concentration. (b) Normalized concentration parameters: High and low frequency susceptibility χhf and χlf , magnetization of saturation M_s , and magnetization remanent of saturation M_{rs} .

sampled profile.

5. Discussion and conclusion

Based on data from a previous analysis, the magnetic concentration parameters $(M_{rs}, M_s \text{ and } \chi)$ reveal that there are different zones that might be related with climate changes. In fluvial deposits the magnetic signal and climate are very likely to be associated through environmental changes that modify the fluvial transport (Jordanova et al., 1997). In flood basin deposits, fine-grained sediments include detritic magnetic grains transported as suspension material, with a silt-clay layer 1 or 2 cm in thickness usually being deposited during one flood period (Reinick and Sing, 1975). The magnetic concentration in the lower part of the LM section (-0.78 to -1.49 m) was found to be higher than that in the upper part (-0.54 m to -0.779 m), indicating that the higher concentration of magnetic minerals in the latter might be related with past humid periods when there was a higher frequency of flooding. The concentration peaks of the four grain types observed in Fig. 6(a) may be related with variations that correspond to four flooding events of different intensity. In this context, diverse lines of paleoenvironmental investigations suggest that during the Holocene the Chaco-Pampean region was characterized by increases in both temperature and precipitation, with a few general departures from these general trends. The period spanning \sim 8.5–3.5 kybp consisted primarily of a humid subtropical and tropical climate, with Brazilian fauna characterized by pedogenesis and fluvial dynamics (Iriondo and García, 1993). Although there is no consensus on the subject of climate (see Tonni et al., 1999), between \sim 7.5– 5 kybp the spectrum of mammals reflect a humid climate (Prado and Alberdi, 1999). Most scholars currently agree that there was a climate change in Buenos Aires province during 5 kybp (Iriondo and García, 1993; Iriondo, 1995; Prieto, 1996; Prado and Alberdi, 1999; among others) and that a humid regime may have existed at the time of the Holocene marine transgression during a period of high sea

level at \sim 5–6 kybp (Isla, 1989). This phenomenon has been reported to have occurred in several places along the cost of the Buenos Aires province (see Weiler, 1994; Pirazzoli, 1996).

In the ancient fluvial system of Buenos Aires city and the surrounding areas, LM is located in the basin of the Maldonado creek, the main water course that exists in the region. Maldonado creek and the Riachuelo-Matanzas River were two significant affluents of the Río de la Plata (Castellanos, 1975), which is a great mass of water that has been suffering significant variations in its geomorphologic configuration since the Late Pleistocene and Holocene. A regional study performed at the Río de la Plata obtained a curve of relative sea level variation during the Holocene that indicated that these changes have produced important modifications in the distribution of submerged and emerged lands (Cavalotto et al., 1995). Based on these results, Cavallotto et al. (1995) concluded that (1) the maximum high stand of sea level at 6 kybp was at an altitude of +6.5 m; (2) between 5–3.5 kybp there was a period of stability with no major variation at +5 m; (3) between 3.5 and 2.9 kybp the sea level rapidly fall to 2.5 m. The Río de la Plata was therefore higher during the middle Holocene than it is at the present time, and a more dynamic drainage system was present at LM. In support of this hypothesis, a marine ingression was dated at \sim 4–5 kybp at about 10 km south of this site in the Riachuelo-Matanzas river basin at Ezeiza (Di Micco, 1990).

In summary, the LM section may be considered to be a soil deposit formed during the Holocene. The magnetic signal showed changes that are to a certain degree correlated with the variations in geology and climate that occurred during the middle Holocene. Consequently, this signal any reflect the humid period prevailing in the region and be representative of the environmental variability existing in northeastern Buenos Aires province during that period. We have shown here that environmental magnetic research is a useful analytical tool for studying issues on climate variability that occurred in the Chaco-Pampean region during more recent geological times.

Acknowledgments. We are indebted to: the University of Buenos Aires and CONICET for their support; Marcia Ernesto (IAG, University of Sao Paulo) for allowing the authors to use their facilities and instruments; the NSF and IAI program for Latin American Quaternary research to global change studies, for providing the AMS dating. The AMS measurement and age calculations were performed by the NOSAMS facility at the Woods Hole Oceanographic Institute and the CU-Boulder INSTAAR Laboratory for AMS Radiocarbon; all other preparations of the samples were performed by the CU-Boulder INSTAAR Laboratory for AMS Radiocarbon Preparation and Research (University of Colorado at Boulder). H. Morrás and C. Favier-Duvois kindly observed the LM soil section; M. Cuadrado Woroszylo was very helpful during the sampling and the editing of this paper.

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