

# Constrained simultaneous algebraic reconstruction technique (C-SART) —a new and simple algorithm applied to ionospheric tomography

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A simple and relatively fast method (C-SART) is presented for tomographic reconstruction of the electron density distribution in the ionosphere using smooth fields. Since it does not use matrix algebra, it can be implemented in a low-level programming language, which speeds up applications significantly. Compared with traditional simultaneous algebraic reconstruction, this method facilitates both estimation of instrumental offsets and consideration of physical principles (expressed in the form of finite differences). Testing using a 2D scenario and an artificial data set showed that C-SART can be used for radio tomographic reconstruction of the electron density distribution in the ionosphere using data collected by global navigation satellite system ground receivers and low Earth orbiting satellites. Its convergence speed is significantly higher than that of classical SART, but it needs to be speeded up by a factor of 100 or more to enable it to be used for (near) real-time 3D tomographic reconstruction of the ionosphere.

**Key words:** Ionosphere, TEC, tomography, GNSS, SART, differential code biases.

## 1. Introduction

Computerized tomography (CT), developed in the 1960s, continues to play an important role in the field of medical imaging. The algebraic reconstruction technique (ART), the first algorithm used for CT (Gordon *et al.*, 1970), is poorly suited for real-time tomographic applications because its iteration steps are time-consuming.

The simultaneous algebraic reconstruction technique (SART), a refinement of ART developed by Andersen and Kak (1984) that solves multiple equations simultaneously, is better suited for real-time applications. It is used in radiological and medical applications, seismic investigations, material science, among others. From a mathematical point of view, the applications differ greatly. In medical applications, the number of observations  $M$  used to reconstruct an image exceeds or is close to the number of unknowns  $N$  (i.e., pixel or voxel values), whereas, in most geophysical applications,  $M \ll N$  is valid (Ivansson, 1986). A good overview of tomographic applications was given by Raymund (1995), who focused on the reconstruction of the ionosphere in detail. Computerized ionospheric tomography (CIT) as a dedicated application of CT has attracted the interest of the scientific community since the navigation satellite systems allows to the derivation of compute ionosphere propagation characteristics. A variety of imaging strategies have been developed within the last years; these allow estimation of electron density fields and enable the study of temporal and spatial variations of the ionosphere. In order to solve the under-determined inversion

problem, several approaches including regularization techniques (e.g., Lee *et al.*, 2007), neural network methods (Ma *et al.*, 2005), Kalman filters (e.g., Hernandez-Pajares *et al.*, 1999; Ruffini *et al.*, 1998), singular value decomposition (e.g., Bhuyan *et al.*, 2004), consideration of background models (e.g., Spencer *et al.*, 2004) and improvements of the SART method (e.g. Wen, 2007; Wen *et al.*, 2007a, b, c), have been developed. Although all of these techniques can reconstruct the probed media with high accuracy, many of them strongly depend on matrix operations, which increases the computation load significantly when the number of unknowns,  $N$ , is large.

We have extended SART, which does not depend on matrix operations, to enable it to carry out tomographic inversions accurately using simple physical relationships.

## 2. Simultaneous Algebraic Reconstruction Technique (SART)

A linear imaging problem such as tomography can be expressed as

$$\mathbf{b} = \mathbf{A} \mathbf{x}, \quad (1)$$

where  $\mathbf{b}$  represents observations  $(b_1, b_2, \dots, b_M)^T (\in \mathbb{R}^M)$ ,  $\mathbf{A} (= (A_{i,j}))$  represents an  $M \times N$  matrix,  $\mathbf{x} (= (x_1, x_2, \dots, x_N)^T \in \mathbb{R}^N)$  stands for the unknowns, and  $T$  is the transpose operator acting on a vector or matrix. SART, as described by Andersen and Kak (1984), is given by

$$x_j^{(k+1)} = x_j^{(k)} + \frac{\omega}{A_{\oplus,j}} \sum_{i=1}^M \frac{A_{i,j}}{A_{i,\oplus}} (b_i - \bar{b}_i(x^{(k)})) \quad (2)$$

for iterations  $k = 0, 1, \dots, K$ . We set the relaxation parameter,  $0 < \omega < 2$ , to 1 for our study. Although larger values speed up convergence, if the value is too large, too much

weight is given to the last projection, which prevents convergence. Smaller values (close to zero) cause the algorithm to converge slowly, which is unsatisfactory for real-time applications and systems with a huge number of cells. Wen *et al.* (2007b) presented an improved algebraic reconstruction technique (IART) based on classical ART. It computes the relaxation parameter at each iteration step adaptively. As the underlying mathematical statistics and the prerequisites for the unknowns remain unstudied, this improved ART is not taken into account here because it is not clear how the introduction of constraints (Section 3) affects IART.

Two definitions are needed for the calculation of expression (2).

$$A_{i,\oplus} = \sum_{j=1}^N A_{i,j} \quad \text{for } i = 1, 2, \dots, M \quad (3)$$

$$A_{\oplus,j} = \sum_{i=1}^M A_{i,j} \quad \text{for } j = 1, 2, \dots, N \quad (4)$$

$$\bar{\mathbf{b}}(\mathbf{x}^{(k)}) = \mathbf{A} \mathbf{x}^{(k)} \quad (5)$$

In classical ART, two prerequisites have to be fulfilled.

$$A_{i,j} \geq 0 \quad \text{for } i = 1, 2, \dots, M \text{ and } j = 1, 2, \dots, N \quad (6)$$

$$\left. \begin{array}{l} A_{i,\oplus} \neq 0 \\ A_{\oplus,j} \neq 0 \end{array} \right\} \text{ for } i = 1, 2, \dots, M \text{ and } j = 1, 2, \dots, N \quad (7)$$

Jiang and Wang (2003) showed that, for  $k \rightarrow \infty$ , Eq. (2) converges to a solution for expression (1) and proved that the result obtained is equivalent to a weighted least squares solution of Eq. (1). For  $M < N$ , the matrix used in the classical least squares adjustment (Koch, 1988) is a singular type and thus cannot be inverted. Singular value decomposition or regularization (Hansen, 1987) is used in this case to obtain a solution for expression (1). Although SART always iterates towards a unique solution independent of  $M > N$  or  $M < N$ , the physical meaning of the results is not given for most under-determined systems.

Equation (1) can be related to tomography applications by denoting the value of cell  $j$  as  $x_j$ . Furthermore,  $A_{i,j}$  can be understood as the length of ray  $i$  in the  $j$ -th cell. Thus, the quantity  $A_{i,\oplus}$  is equal to the total length of the  $i$ -th ray, and  $A_{\oplus,j}$  is the sum of all ray paths crossing the  $j$ -th cell. Since the ray length is always a positive number, Eq. (6) and the first case of condition (7) are fulfilled. If cells  $t_\alpha$  ( $\alpha = 1, 2, \dots, A \leq N$ ) are not crossed by any ray (i.e.,  $A_{\oplus,t_\alpha} = 0$ ), division by zero would occur in Eq. (2). This problem can be easily solved by applying the algorithm to only cells that are traversed by at least one ray—i.e., ( $j \mid j \in \{(1, 2, \dots, N) \wedge j \notin t_\alpha\}$ ).

### 2.1 Applying SART to GNSS ionosphere tomography

To reconstruct the electron density distribution of the ionosphere using data from the global navigational satellite system (GNSS), one has to take into account that satellite and receiver effects bias the data. Thus, the observation equation obtained using dual-frequency code measurements or L1–L2 leveled phase measurements (both described, for example, by Schaer (1999)) basically reads as

$$\text{STEC}_{\text{obs}} = \text{STEC} + \text{DCB}^s + \text{DCB}^r, \quad (8)$$

where STEC is the slant total electron content measured in total electron content units ( $1 \text{ TECU} = 10^{16} \text{ electrons m}^{-2}$ ). Differential code biases (DCBs) are assigned to both the satellite ( $s$ ) and receiver ( $r$ ) offsets (e.g., Ray and Senior, 2005) and are added to the slant total electron content so that  $\text{STEC}_{\text{obs}}$  is obtained from the raw data (Eq. (8)). Moreover, for ray  $i$ ,

$$\text{STEC}_i = \sum N_e(j) \Delta s_{i,j}, \quad (9)$$

where  $j$  denotes cells crossed by the ray,  $N_e(j)$  represents the electron density of cell  $j$ , and  $\Delta s_{i,j}$  is the path length inside the cell. It is obvious from this notation that  $N_e(j) \equiv x_j$  and  $\Delta s_{i,j} \equiv A_{i,j}$ . Since it is not possible to estimate satellite and receiver DCBs together without setting a reference level, it is common (Schaer, 1999) to place a zero-sum condition on the satellite biases, i.e.,  $\sum_{s=1}^S \text{DCB}^s = 0$ , where  $S$  denotes the number of satellites. To estimate the satellite and receiver DCBs when using SART, one has to treat them as “artificial” cells, with the exception that path length  $A_{i,j}$  is always equal to one. Additionally, one has to place the same condition on the satellite DCBs. Thus,  $b_i = 0$  ( $i$  being the number of artificial observations) has to be applied in order to set up the zero-sum condition when using SART.

### 3. C-SART—an Extension of SART for the Reconstruction of Smooth Fields

As described above, SART has the advantage, compared to least-squares adjustment, that even high-resolution tomographic problems, which entail a huge number of unknowns, can be solved without reaching the limits of computing power and memory. The number of mathematical operations in a tomographic problem solved using SART scales is determined by the number of unknowns ( $N$ ), whereas least squares adjustment, Kalman filter (Kalman, 1960) methods, and the singular value decomposition scale are determined mainly by the size of the design matrix, which is  $N \times N$ . For example, the computation time for matrix inversion follows  $\sim O(N^3)$ , which makes SART more efficient than the three approaches mentioned above, even though it is an iterative technique.

The Kalman filter approach has been used in several studies to obtain highly resolved images of the ionosphere (e.g., Hernandez-Pajares *et al.*, 1999). With this approach, the physical conditions of the media can be flexibly added as additional information, which supports the observation geometry in cells for which no information was gathered (Hajj *et al.*, 2004). Ideally, the number of rays in tomography is larger than the number of cells, so the ray geometry defines a non-singular matrix, which permits reconstruction of the probed media. This is not usually the case in geoscience applications, unlike medical applications. To handle such situations, we applied a new approach to SART by using a simple physical model that supports the estimation of the electron density field. We term our version of SART “constrained-SART (C-SART)”. In the following, we will use the expression “constraint” for any kind of artificial observation which supports the solution of the tomography problem.

### 3.1 Constrained-SART: basic idea

Figure 1 illustrates the basic idea of C-SART for a two-dimensional case (2D) in which there are nine cells (with values  $x_j$ ,  $j = 1, 2, \dots, 9$ ), and a ray never crosses the middle cell ( $x_5$ ). In this case, information about the value of the center cell can be transferred from the neighboring cells if a number of assumptions are made about the underlying field. Generally, any physically reasonable information, expressed in the form of a finite difference equation, can be transferred. Treating the media as a smooth field—i.e., no steep gradients between neighboring cells—is a reasonable way to simplify the discussion and to enable this approach to be extended to applications other than ionosphere tomography. One way to do this is to use the 2D-Laplacian operator,

$$\hat{L} = \begin{bmatrix} -1 & -1 & -1 \\ -1 & 8 & -1 \\ -1 & -1 & -1 \end{bmatrix}, \quad (10)$$

which relates the values for the neighboring cells to that of the cell at the center. This operator is used to compute the difference between the sum of the values for the surrounding cells and the value for the center cell, which is multiplied by the number of neighboring cells. Since the mean value of all entries in the matrix is zero, application to the ionosphere does not bias the total number of electrons in the field. The operator has to be modified accordingly when the concerned cell lies on the edge of the model space. It can be used to introduce an artificial observation between unknowns, a so-called “constraint”. Thus, it is possible to impose, for the nine-cell example, one constraint for each by applying the smoothness operator to the surrounding cells. This simple approach can be applied to any cell, independent of the number of crossing rays, to guarantee that the estimated field is smooth. This constraint, expressed as a function of  $f_c$  in its general form, can be denoted as

$$\frac{1}{\beta} \cdot f_c(x_a, \dots, x_b) = 0, \quad (11)$$

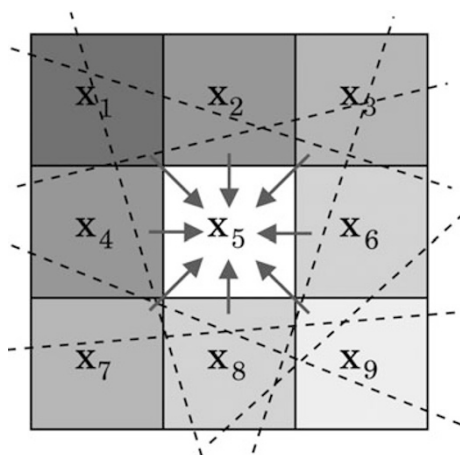


Fig. 1. Basic idea of C-SART for two dimensional (2D) case. The  $x_j$  ( $j = 1, 2, \dots, 9$ ) represent the cells, which are gray shaded on an arbitrary scale. Cell at center ( $x_5$ ) is never crossed by a ray. Placement of a Laplacian constraint on the underlying field enables the value of  $x_5$  to be deduced from those of neighbor cells.

where  $x_a, \dots, x_b$  includes all of the unknowns related to the condition. Each constraint is weighted by hyper-parameter (weight parameter)  $\beta$ , which determines the extent to which the condition must be fulfilled. The choice of  $\beta$  should be cross-validated to ensure that the reconstructed field is neither too rough nor too smooth. Expressing the Laplace operator as a constraint yields

$$\frac{1}{\beta_L} \left( Cx_j - \sum_{c=1}^C x_c \right) = 0, \quad (12)$$

where  $C$  is the number of neighboring cells, which can range between 3 and 8 (for the 2D case) depending on the position of cell  $j$  in the grid. It is thus possible to set up one constraint for each cell that increases the number of observations (by the number of cells) and makes the whole system overdetermined. For small grids, the solution can be easily obtained using traditional least-squares adjustment since redundancy is ensured by the constraints. Even for a 2D case of  $100 \times 100$  cells, SART is much faster than the inversion of the corresponding  $10,000 \times 10,000$  design matrix from the Gauss-Markov model. In general, heat-, wave- or Laplace-partial derivative equations can be expressed using finite differences, and dedicated constraint operators, similar to expression 10, can be set up. Moreover, it is possible to support the estimates by considering physical relationships given as explicit equations. For the case of the ionosphere, it might be useful to constrain the solution to follow Chapman-like vertical profiles (as described, for example, by Hargreaves, 1992). Thus, it is possible to write

$$x_j - x_m \exp \left[ 1 - \left( \frac{h(x_j) - h_m}{H} \right) - \exp \left( -\frac{h(x_j) - h_m}{H} \right) \right] = 0, \quad (13)$$

where  $x_m$  is the electron density maximum and  $h_m$  the corresponding height for each vertical profile. Operator  $h(x_j)$  returns the height of cell  $j$ , and  $H$  is the scale height of a hydrostatic equilibrium (e.g., Hargreaves, 1992). To apply this constraint within C-SART, it is necessary to compute  $x_m$  and  $h_m$  from the results of the prior iteration step. In general, the Laplacian constraint could be used together with the Chapman profile approach if the user wants to force the resulting field to follow simple physical conditions of the ionosphere. For the tomographic inversion discussed in Section 4.2, such a Chapman constraint was not applied since our intention was to demonstrate that C-SART performs well even without knowledge of the underlying field. As discussed above, for the case in which DCBs have to be estimated together with the electron density field, it is common to constrain the sum of the satellite DCBs to zero to achieve a virtual but stable reference (Schaer, 1999). One can express this zero-sum condition as

$$\frac{1}{\beta_D} \sum_{d=1}^D x_d = 0, \quad (14)$$

where  $x_d$  are the cells corresponding to the DCBs, and  $\beta_D$  is the corresponding weighting factor.

### 3.2 Mathematical prerequisites

Equation (12) violates condition (6) since  $A_{i,j} < 0$  is possible. Moreover, when the coefficients of the imaging system take negative values,  $A_{i,\oplus}$  is negative. Therefore, the original SART had to be refined. Censor and Elfving (2002) showed that SART convergence is ensured when

$$A_{i,\oplus} = \sum_{j=1}^N |A_{i,j}| \quad \text{for } i = 1, 2, \dots, M \quad (15)$$

and

$$A_{\oplus,j} = \sum_{i=1}^M |A_{i,j}| \quad \text{for } j = 1, 2, \dots, N \quad (16)$$

are used instead of Eqs. (3) and (4). Expression (2) does not need to be changed. The use of Eqs. (15) and (16) to compute  $A_{i,\oplus}$  and  $A_{\oplus,j}$  enables the C-SART algorithm to be handled with the SART formalism described by Eq. (2).

## 4. 2D Reconstruction of Ionosphere Using C-SART—A Test Case Using Artificial Data

We used a 2D test scenario for testing the performance of the C-SART algorithm in comparison with the SART algorithm. One hundred ionosphere profiles were computed using the international reference ionosphere (IRI) model (Bilitza, 2001), version IRI2007. Data were obtained for the Greenwich meridian for latitudes between  $-49^\circ$  and  $50^\circ$  for May 6, 2004, 1200 local time. The height ranged from 105 to 600 km in steps of 5 km, resulting in a grid of  $100 \times 100$  electron density values (Fig. 2). This 2D electron density field was used as a reference for our investigation characterizing the quality of the tomographic inversion. Artificial observations were assumed to be inside the plane of the profiles only, enabling us to treat the problem as a 2D one. Additionally, the curvature of the Earth was ignored to simplify the discussion. It was also assumed that the dispersive delays outside the model boundaries had been removed from each observation; this can be achieved with the help of theoretical models or simple approximations of the plasmasphere contribution (e.g., Ma *et al.*, 2005). For investigations including the plasmasphere and higher altitudes, it would be worthwhile adding a coarse voxel structure above the ionosphere domain. The electron densities of these voxels can be obtained together with the ionosphere-related values from the same tomographic inversion.

To obtain a realistic, but weak spatial distribution of the receivers, we assumed that the ionosphere was probed by 14 ground receivers and two low Earth-orbiting (LEO) satellites, which provided occultation data. The number of GNSS satellites traceable by all receivers was set to six. Data from three arbitrary epochs were used to reconstruct the media (Fig. 3). This test geometry does not totally reflect a real-world scenario since the GPS-to-ground geometry changes much less frequently than LEO occultations occur. Nevertheless, since our investigation focused on the improvement in the reconstructed electron density field due to usage of C-SART, we can draw conclusions from our results about how GNSS applications can benefit from C-SART. Once the algorithm has been implemented in a way that permits fast computation of dense 3D electron density

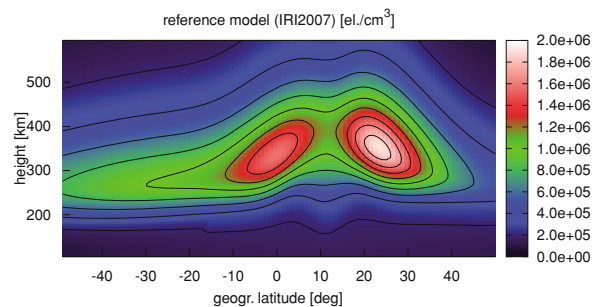


Fig. 2. Reference ionosphere latitude profile generated from the IRI2007 model run assuming May 6, 2004, 1200 LT, at  $0^\circ$  longitude. Obtained electron densities are referenced to the mid-point of each cell.

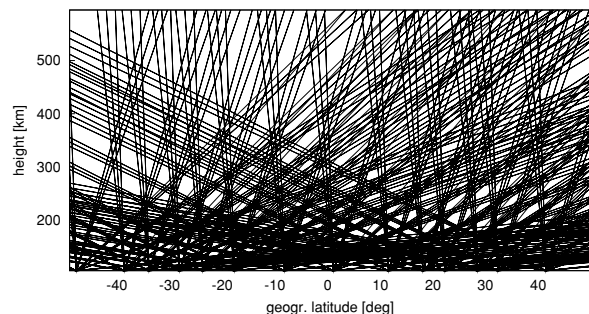


Fig. 3. Ray geometry obtained from 14 ground receivers and two LEO satellites, which provided occultation data. Six GNSS satellites were traceable by all receivers. Each receiver tracked satellites in three different epochs, with angular separation between consecutive epochs of 1 degree.

fields, we plan to test it using data from dense ground GNSS receiver networks.

Slant total electron content values were obtained for each observation by ray-tracing through the reference ionosphere, ignoring contributions from electrons at higher altitudes or outside the latitude boundaries. A total of 10,022 unknowns (including electron density values for the cells and the receiver and satellite DCBs) were estimated from 288 ( $= (14 + 2) \cdot 6 \cdot 3$ ) observations, which is a highly under-determined situation for tomographic inversion. Of the pixels, 11% were not hit by any ray, and 21.9% were hit by only one ray. Therefore, with traditional SART, many cells would need support from background models or would even be removed from the tomographic inversion. In total, about one-third of the field would lack good ray coverage and thus would not be reconstructed unbiased.

The DCB values for the receivers and satellites were added to the ray-traced STEC values to simulate actual GNSS conditions. The ambiguities were assumed to have already been resolved (e.g., Horvath and Crozie, 2007), so it was possible to treat the data like code-leveled phase measurements. The artificial measurements were corrupted with Gaussian random noise, at a signal-to-noise ratio (SNR) of 100, to obtain more realistic signal characteristics.

### 4.1 Optimum choice of constraint weights by model verification

Since the “real” electron density field is known, the tomographic reconstruction results could be easily cross-

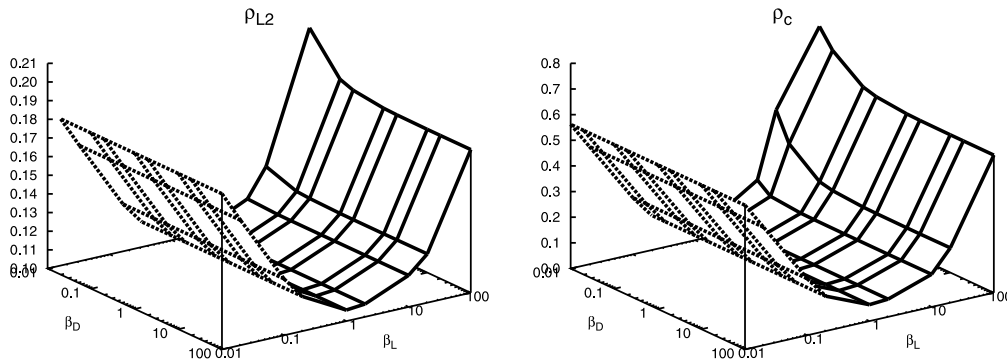


Fig. 4. Results of cross-validation test after 1,000,000 iterations for different values of constraint weights  $\beta_L$  and  $\beta_D$ . Left plot shows results for L2-space metric,  $\rho_{L2}$ ; right plot shows them for c-space metric,  $\rho_c$ .

validated. Andreeva *et al.* (1992) proposed two measures for describing the deviation between two models. In one, the metric is assigned to the L2 space; in the other, it is assigned to the c-space. The first metric,

$$\rho_{L2} = \sqrt{\frac{\sum_{i=1}^N (\tilde{x}_i - x_i)^2}{\sum_{i=1}^N \tilde{x}_i^2}}, \quad (17)$$

is the ratio between the standard deviations of the differences and the original field values  $\tilde{x}_i$ . Thus, a small value of  $\rho_{L2}$  can be taken as an indicator of good global performance for tomographic inversion. The second metric,

$$\rho_c = \frac{\max_i |\tilde{x}_i - x_i|}{\max_i |\tilde{x}_i|}, \quad (18)$$

utilizes local performance characteristics by relating the largest reconstruction error to the largest true value. The better the performance, the smaller the metric. Thus, to find the optimum constraint weight parameters  $\beta_L$  and  $\beta_D$ , we computed  $\rho_{L2}$  and  $\rho_c$  for different weights. Figure 4 shows the results of the cross-validation test after one million iterations for different values of the constraint weights. Agreement with the model strongly depended on the value set for  $\beta_L$ , and agreement was best when  $\beta_L$  was set to one. The selection of the  $\beta_D$  value was less critical as it did not have a noticeable effect on  $\rho_{L2}$  and  $\rho_c$  as long as  $\beta_L > 0.1$ . Since  $\beta_L$  strongly determines the roughness of the reconstructed field, setting of its value can lead to bigger reconstruction errors when the field is forced to be too smooth ( $\beta_L < 1$ ). If  $\beta_L$  is set to a value larger than one, the algorithm cares less about the Laplacian constraints and tends to perform in a way similar to classical SART. Thus, in the following examples,  $\beta_L = \beta_D = 1$  is used for the reconstruction of the electron density field. For tomographic problems with a larger number of rays, a cross-validation test should be done again to determine an optimum pair of values for  $\beta_L$  and  $\beta_D$ .

#### 4.2 Tomographic reconstruction results for SART and C-SART

Using SART and C-SART with the optimum constraint weights from above, we reconstructed the electron density

Table 1. Modeled and reconstructed (after  $10^6$  iterations) DCB values for SART and C-SART algorithms.  $DCB_s^s$  represents value for GNSS satellite  $s$ , and  $DCB_r$  represents value for receiver  $r$ . Values for receivers on-board LEO satellites are denoted by  $A$  and  $B$ .

Type	Model	SART	C-SART
$DCB^1$	-2.90	-0.24	-2.89
$DCB^2$	0.45	0.13	0.46
$DCB^3$	1.45	0.13	1.55
$DCB^4$	2.10	0.16	2.20
$DCB^5$	0.95	1.40	0.83
$DCB^6$	-2.05	-1.58	-2.15
$DCB_1$	0.50	-0.13	0.61
$DCB_2$	1.00	0.16	1.17
$DCB_3$	1.50	0.18	1.65
$DCB_4$	2.00	0.20	2.09
$DCB_5$	2.50	0.23	2.55
$DCB_6$	0.50	0.16	0.47
$DCB_7$	1.00	0.22	0.86
$DCB_8$	1.50	0.21	1.41
$DCB_9$	2.00	0.23	1.98
$DCB_{10}$	2.50	0.27	2.50
$DCB_{11}$	0.50	0.16	0.56
$DCB_{12}$	1.00	0.19	1.09
$DCB_{13}$	1.50	0.22	1.64
$DCB_{14}$	2.00	0.25	2.14
$DCB_A$	-0.50	0.44	-0.50
$DCB_B$	1.00	0.15	1.07

fields. One million iterations were carried out, with  $\beta_D$  set the same for both algorithms. Figure 5 shows the reconstructed fields for both algorithms. The classical SART algorithm did not reconstruct the model ionosphere well and even produced some negative electron density values. Since all unknowns were initialized with zero values, the cells not crossed by rays retained this value through all iterations. Moreover, the DCBs (Table 1) were not recovered at all, which directly translated into artifacts in the reconstructed image. In contrast, the C-SART algorithm reconstructed the model ionosphere much better. It did not produce negative values, the recovered field looked very similar to the model one (Fig. 1), and the uncrossed cells were updated with information from the neighboring ones by the Laplacian constraint as expected. The absolute relative error of the C-SART reconstruction, as depicted in Fig. 6, did not exceed 75% and was less than 15% for most of the regions. Closer examination revealed that the areas with higher relative errors had lower electron densities, meaning that their absolute reconstruction error was not necessarily large.

Averaging the absolute relative errors over the whole im-

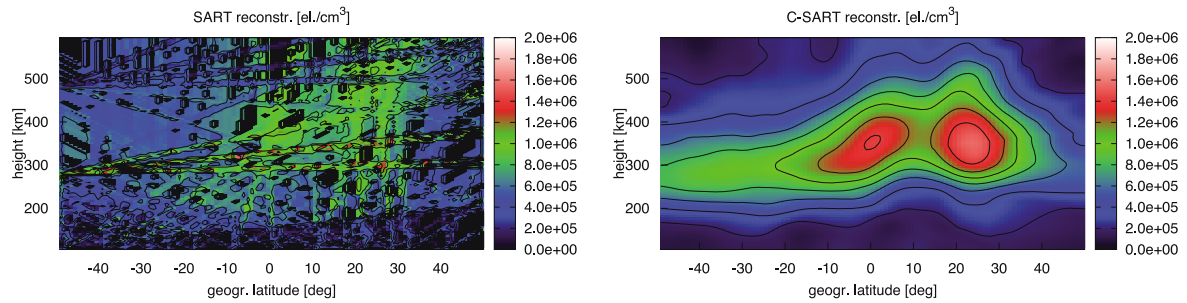


Fig. 5. Reconstructed electron density fields after 1,000,000 iterations. Left figure shows estimated field for SART, and right one shows it for C-SART. Note that the range of the values differs between plots and that color coding is not the same.

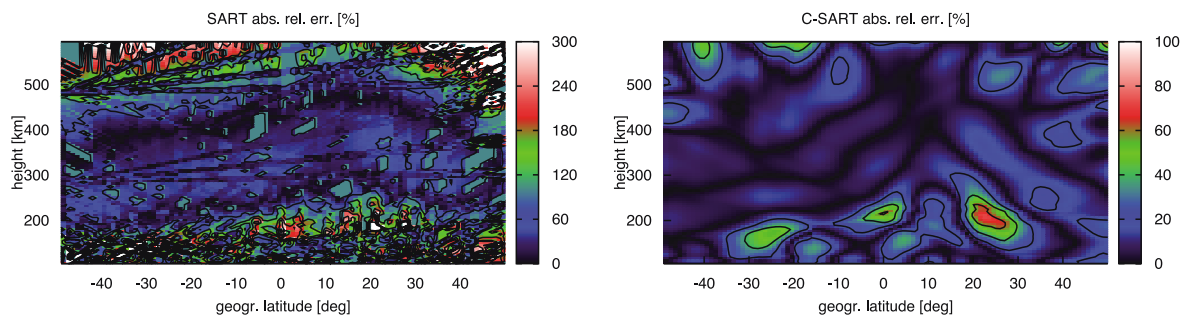


Fig. 6. Absolute relative reconstruction errors with respect to IRI model for SART (left) and C-SART (right) algorithms. Note that the range of the values differs between plots and that color coding is not the same.

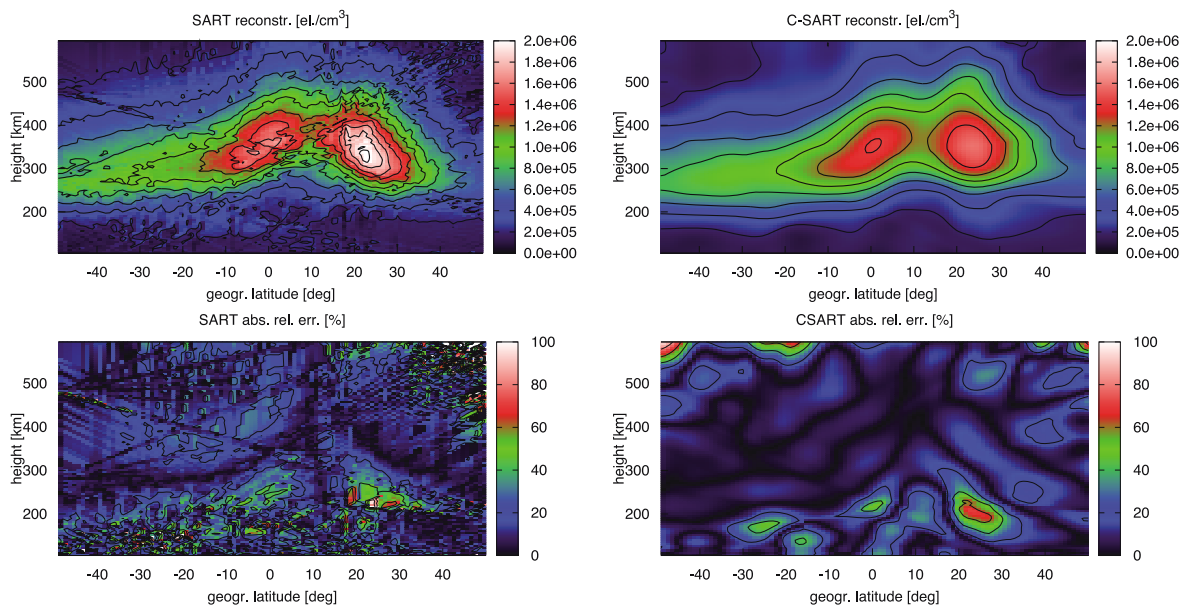


Fig. 7. Electron density fields reconstructed by SART and C-SART when DCBs were known and models were initialized with IRI2007 data from an epoch 1 month earlier. Upper plots show reconstructed fields, and lower plots show corresponding absolute relative errors with respect to original field.

age produced a mean reconstruction error of 11.9%, which is in good agreement with  $\rho_{L2}$ , a similar measure. The absolute relative errors for the SART reconstruction (Fig. 6) were large (82% on average). This clearly demonstrates that C-SART provides much better results than SART.

**4.2.1 Reconstruction of differential code biases**  
Satellite DCBs are usually known up to a certain accuracy level since they are monitored on a daily base by several GNSS analysis centers (Feltens, 2003). The receiver biases

are usually available for download if the station belongs to the global GNSS network. Otherwise they have to be determined in a prior step or be estimated together with the ionospheric parameters. To demonstrate that C-SART performs well even when the receiver and satellite DCBs are unknown, the DCBs were included in the algorithm and were treated as “virtual” cells with the mathematical expressions described in Section 2. Table 1 shows the DCBs recovered with the SART and C-SART algorithms after one

million iterations. C-SART clearly reconstructed the satellite and receiver biases, whereas classical SART did not. The zero-sum constraint, applied to the satellite DCBs, was met in both cases, but only C-SART with its simple Laplacian constraints on the electron density field produced the values correctly. Since SART did not iterate towards the correct satellite DCBs, the receiver DCBs and the electron density field itself were negatively affected. The satellite DCBs from C-SART were within 0.17 TECU (total electron content) of the model values. Those for the ground-based and LEO on-board receivers were within about the same range. These estimated DCB values should satisfy user needs, especially when the low number of input observations and the weak observation geometry are considered. The values from SART did not agree with the model ones, negatively affecting reconstruction of the electron density field.

**4.2.2 Known DCBs and initialization with background model** To demonstrate how SART and C-SART perform when all DCBs are known and the models are initialized with a background ionosphere model, IRI2007 electron density profiles were computed for an epoch of 1 month preceding the one used in the previous section. This was done to ensure that the cells were initialized with values that were realistic but also sufficiently different considering that the background model has a limited ability to predict actual conditions. As the receiver and satellite DCBs were known, only the electron density values for each cell had to be estimated. Nevertheless, SART did not update the uncrossed cells, and the values of the background model remained unchanged. Thus, the quality of the background model strongly determines the accuracy with which SART can recover the ionosphere when the geometrical coverage is poor.

The electron density fields reconstructed by the SART and C-SART algorithms are shown in Fig. 7 together with the absolute relative errors. The performance of SART was greatly improved: the average absolute reconstruction error was 16.1%, and the maximum was 235%. The C-SART algorithm benefited only slightly from the background model initialization: the mean absolute relative reconstruction error improved slightly to 11.6%. However, the field reconstructed by C-SART was slightly degraded in the upper ionosphere (the maximum error in that region was 98.1%). Therefore, the use of a background model is significantly useful for the SART algorithm, but using it still does not solve the problem of uncrossed cells. The C-SART algorithm performs about 50% better than the SART one regardless of whether there is knowledge of the DCBs or the background model is initialized. The only advantage gained from using the background model for the C-SART algorithm is that the convergence speed is slightly faster, as discussed in the next section.

**4.2.3 Convergence and computation speed** The sum of the squared improvements (SSI) for iteration  $k$ ,

$$\text{SSI} = \sum_{j=1}^N \left[ \frac{\omega}{A_{\oplus,j}} \sum_{i=1}^M \frac{A_{i,j}}{A_{i,\oplus}} (b_i - \bar{b}_i(x^{(k)})) \right]^2, \quad (19)$$

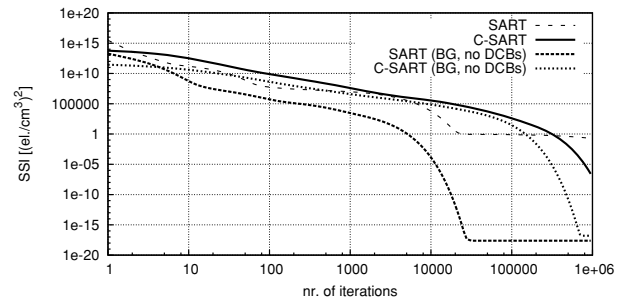


Fig. 8. Convergence of SART (dashed line) and C-SART (thick line) measured using sum of squared improvements for each iteration step. The thick dotted lines correspond to SART and C-SART runs with known DCBs and background field initialization.

was used as an indicator of the convergence speed. Equation (19) is also related to the energy (Mailloux *et al.*, 1993) remaining in the system after the  $k$ -th iteration. Thus, it can be used to define a threshold for stopping the iterative process if SSI decreases to a certain value. Figure 8 plots SSI against the number of iterations for the SART and C-SART algorithms with and without DCB information and background model initialization. When the DCBs were known and the background model was initialized, SSI stabilized before one million iterations for both algorithms. However, when the DCBs were estimated and the models were initialized with zero values, the SSI for SART did not decrease much, and that for C-SART saturated only after about 1.2 million iterations (not shown here). For practical applications, it is sufficient to stop the iterations once SSI drops below  $10^{-2}$  (electrons<sup>2</sup>/cm<sup>6</sup>) as any subsequent improvements are usually too small to affect the tomography results. It took 737 s to complete  $10^6$  iterations with SART and 1418 s with C-SART. C-SART reached the threshold of  $10^{-2}$  (electrons<sup>2</sup>/cm<sup>6</sup>) in about  $2 \cdot 10^5$  iterations. Thus, it took only 4 min and 44 s (i.e., 284 s) on a simple PC (2.3-GHz Pentium D CPU, 2-GB RAM) to carry out high-resolution ionospheric tomography and estimate the unknown DCB values (although their values might have stabilized in an earlier iteration). The coding of the C-SART algorithm in a low-level programming language will lead a significant reduction in computing time (by at least an order of magnitude).

## 5. Discussion

Our simple algorithm, the constrained simultaneous algebraic reconstruction technique (C-SART), has many advantages for reconstructing the ionosphere by means of electron density values using GNSS measurements. It is suitable for real-time applications and even enables estimation of instrumental biases within a reasonable processing time. The reconstructed fields are significantly better than those obtained from classic SART. Moreover, C-SART can estimate the DCB values accurately enough when necessary.

In the test case we used, in which a very weak sampling of the 2D field was assumed, the average absolute error of the reconstructed field was about 12% (including cells not crossed by rays). The estimated DCB values had a maximum error of about 0.2 TECU, which is accurate enough

for an unbiased reconstruction of the ionosphere. Since C-SART is based only on the assumption that the underlying field is smooth, it can be applied to other tomography applications as well. It is also easy to apply other conditions in the form of difference equations to the algorithm, e.g., heat-, wave-, and Laplace-partial derivative equations. This means that physical conditions estimated from measurement data can be considered in the field reconstruction, resulting in better representation (including unprobed areas) of the media. One drawback of the smoothness operator is that, for applications in which the media contain discontinuities, the C-SART algorithm gives incorrect results. In the case of steep gradients, the grid should be refined or  $\beta_L$  should be increased in order to allow for more variation between cells.

## 6. Future Work

To enable the C-SART algorithm to be used for (near) real-time 3D tomography of the ionosphere, we need to speed it up by a factor of 100 or more. Assuming that GNSS observations are taken every 30 s and that some time is spent on data transfer and the leveling of the L1–L2 phase measurements, about 15 s should remain for ionospheric reconstruction. A speed-up by a factor of 10 or can be achieved by optimizing the routines and loops and by coding parts of the algorithm in a low-level programming language (e.g., Assembler). In addition, the reconstruction could be modularized using either message parsing interfaces (described, for example, by Skjellum *et al.*, 1993) or OpenMP™ (Dagum and Menon, 1998), which distributes the computation load among multi-core processors.

The C-SART algorithm can thus be speeded up (nearly) proportional to the number of core processors used. Moreover, Moore's Law (Moore, 1965), which predicts the number of transistors in future CPUs, indicates that the realization of on-line monitoring of the ionosphere using C-SART is feasible. Additionally, a large number of iterations ( $2 \cdot 10^5$ ) is not necessary when models are generated in 30-s intervals since each model run can be initialized with the results of the prior one, which will not differ much from the new one. Moreover, other measurements, such as ionosonde profiles, can be utilized when available.

Although C-SART has been applied to ionosphere tomography it is not necessarily limited to this single application. Basically, it can be used for any kind of tomographic inversion as long as constraints can be defined in a meaningful sense. For example, C-SART can be used for seismic prospecting by applying a-priori velocity information as constraints. It can be used for other problem statements occurring in seismology, atmosphere, space-physics, and medical research without any large modifications. Thus, the flexibility of C-SART, paired with increasing computational power, will make it a powerful tool for a variety of scientific applications.

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