Observations of mesospheric sporadic sodium layers with the MU radar and sodium lidars

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The dynamical structure of the atmosphere around the sporadic sodium layer at mid-latitude (\sim 35°N) below 100 km was studied by simultaneous observation with the MU radar at Shigaraki (34.9°N, 136.1°E), and two Na lidars at Shigaraki and in Hachioji (35.6°N, 139.4°E). In the lidar data, fifteen Na_s (sporadic sodium layer) events were detected. Wind shear, temperature, and stability indices, at around the time and height of Na_s were observed with the MU radar. Strong total wind shear correlated well with Na_s, especially when sporadic Es did not accompany. However, no other clear correlations, such as correlations with temperature etc., were found. The result is similar to the report of the lidar observations in Hawaii during the ALOHA-93 campaign (Qian *et al.*, 1998), and suggests a similar generation mechanism between 20°N and 35°N.

1. Introduction

The vertical profiles of the mesospheric sodium layer, distributed at altitudes between about 80 and 110 km, have been observed for more than 20 years by means of lidar techniques. During these observations, researchers discovered the sudden formation of a dense thin Na layer superposed on variations of mesospheric sodium layers. Such an enhanced layer is called a sporadic sodium layer (Na_s) or a sudden sodium layer. An Na_s is characterized as a thin layer with a full-width at half maximum (FWHM) of 1–2 km, typically, lasting for a few tens of minutes to several hours. The ratio of peak Na density to the normal sodium density ranges typically from three to five, but occasionally is as large as ten (Nagasawa and Abo, 1995).

Clemesha *et al.* (1978) were the first to report observations of Na_s layers which continued for three hours at São Paulo, Brazil (23°S) in the late 1970s. In the 1980s, many Na_s events were observed at high latitudes at Andenes, Norway (69°N) (von Zahn *et al.*, 1987; von Zahn and Hansen, 1988), and Longyearbyen, Svalbard (78°N) (Gardner *et al.*, 1988), as well as at low latitude stations, such as Mauna Kea, Hawaii (20°N) (Kwon *et al.*, 1988) and Arecibo, Puerto Rico (18°N) (Beatty *et al.*, 1989). However, at mid-latitudes, Na_s were not observed despite the observation of normal sodium layers for long periods in Illinois (40°N) (Gardner *et al.*, 1986) and Haute Provence, France (44°N) (Granier and Megie, 1982). In the 1990s, the fairly frequent occurrence of Na_s at mid-

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latitude at Hachioji, Japan (35.6°N, 139.4°E) was reported by Nagasawa and Abo (1995), and the question arose as to the reason for such a big difference in the Na_s occurrence at mid-latitudes.

Thus, Na_s has been observed at low and high latitudes, as well as at limited mid-latitude stations, and various generation mechanisms have been proposed and discussed. As for high latitude observations, strong correlations between Na_s and sporadic E (Es) occurrence have been observed, and mechanisms by which Na_s is generated from the Es layer have been discussed (von Zahn et al., 1987; von Zahn and Hansen, 1988). They pointed out that the atomic Na generation by the high energy auroral particles from upper atmospheric smoke or dust particles concentrated on an Es occurrence is a possible mechanism for these high latitude Na_s. A very good correlation between Es and Na_s was observed on low latitude observations using the incoherent scatter radar technique (Beatty et al., 1989; Kane et al., 1993). Based on such observations, the conversion of Na⁺ ions, Na-bearing molecules, or smoke particles into neutral sodium atoms was discussed (e.g., Hansen and von Zahn, 1990; von Zahn and Murad, 1990). Kirkwood and von Zahn (1991) have suggested that a strong electric field generates Es and Na_s layers in the auroral region. It is noteworthy that Nagasawa and Abo (1995) also pointed out the correlation between Na_s and Es based on the observations at Hachioji at mid-latitude. More recently, Cox and Plane (1998) discussed the possibility of downward motion of Es as a source of Na_s formation from Na⁺ based on the results of laboratory experiments.

The possibility of neutral chemistry was also reported as a candidate for Na_s generation. Zhou *et al.* (1993) argued that a temperature increase due to tides and/or gravity wave activity may produce the Na enhancement, since the sodium abun-

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dance is sensitive to temperature. The deposition of sodium from meteors was also discussed based on observations at Sao Jose dos Campos (23°S, 46°W) (Clemesha *et al.*, 1978, 1980, 1988; Batista *et al.*, 1989). Vertical displacement of the sodium layer to form Na_s has also been considered as a possible cause (Clemesha *et al.*, 1996).

Thus, Na_s phenomena have been observed at different latitudes and various possibile sources have been discussed. However, little work has been done on the Na_s at the midlatitude, which have been frequently observed in Hachioji, Japan (35.6°N) (Nagasawa and Abo, 1995), because of the lack of simultaneous campaign observations such as those ones carried out at high and low latitudes with radars and rocket observations.

The MU radar, an MST (mesosphere-stratosphere-troposphere) radar and an IS (incoherent scatter) radar using VHF frequency, located at Shigaraki (34.9°N, 136.1°W), has been utilized to observe the dynamics of middle and upper atmospheres. Recently, the MU radar was used as a meteor radar (Nakamura *et al.*, 1991), and the observation of the dynamical parameters such as wind and temperature fluctuations around the mesopause region has been carried out (Tsutsumi *et al.*, 1996; Nakamura *et al.*, 1997).

In this paper, the results of simultaneous observations with the MU radar and two sodium lidars, at Hachioji and Shigaraki, are used to elucidate the dynamical structure of the mesopause region on the occurrence of sporadic sodium layers at mid-latitude, in order to discuss the source of the Na_s layers at this latitude. The observation of atmospheric waves in the mesospheric sodium layer with the MU radar and the sodium lidar at Hachioji has been reported by Namboothiri *et al.* (1996), and Kobayashi *et al.* (1999), but in this study we focus only on the Na_s events.

2. Observation

2.1 Meteor echo observation with the MU radar

The MU (middle and upper atmosphere) radar is a monostatic pulse Doppler radar operated at 46.5 MHz VHF, constructed at Shigaraki, Shiga, Japan (34.9°N, 136.1°E) in 1984 (Fukao et al., 1985a, b). The MU radar is characterized by an active phased array system with 475 Yagi antennas, and has been utilized for various middle atmospheric and upper atmospheric studies. This radar has been applied to meteor observations recently, in order to observe the mesosphere and lower thermosphere regions for 24 hours a day (Nakamura et al., 1991). The specific parameters of the meteor echo observation mode are summarized in Table 1. For the meteor mode observation, a transmitted pulse is phase modulated by a 16-bit complementary code with a 6 μ s or 8 μ s subpulse width. The inter-pulse period (IPP) is 1.95–2.56 msec. The received signal is coherently integrated for 4–8 pulses. The decay rate of meteor echo power is calculated to obtain the diffusion coefficient. The doppler frequency shift is also determined from the phase progression of the meteor echo, from which the radial wind is estimated. We detect about 10,000 to 20,000 meteor echoes per day, and the observed radial wind and diffusion coefficient are accumulated in a time-height bin of 30 minutes \times 1 km. Then, the diffusion coefficient and horizontal wind velocities are estimated in each bin (Nakamura et al., 1997).

Table 1. Specifications of the MU radar meteor mode.

Location	Shigaraki (34.9°N, 136.1°E)
Transmitter	Circular array of 475 crossed Yagis
Operational frequency	46.5 MHz
Peak power	1 MW (maximum)
Average power	50 kW
Pulse length	6–8 μsec
Code modulation	16 bit complementary code
Beam	all azimuth, elevation angle $60^{\circ}/45^{\circ}$
Receiver	3 element Yagi × 4 (interferometer)
Range resolution	900 m
Sample range	76 km–248 km
Wind velocity & diffusion coe	efficient data
Height resolution	1 km
Time resolution	30 minutes

2.2 Lidar observations of the mesospheric sodium layer

In this study we employed two sodium density lidars; one at Tokyo Metropolitan University (TMU) in Hachioji (35.6°N, 139.4°E) and the other at the MU radar observatory in Shigaraki (34.9°N, 136.1°E). The former has been developed and operated by TMU and the latter by Shinshu University. The two lidars have a similar system diagram except for the energy source of the dye laser and the tuning system.

The TMU Na lidar system consists of a tunable dye laser, pumped by a frequency doubled Nd:YAG laser, a receiving telescope, a photomultiplier tube detector, a monitoring system for the transmitted laser wavelength, and a data acquisition system (Nagasawa and Abo, 1995). The data acquisition was performed by photon counting, with a sampling gate of 667 ns, corresponding to a range resolution of 100 m. A receiving Cassegrainian telescope with a primary mirror with a diameter of 28 cm (—October, 1992), 35 cm (October, 1992—December, 1993), or 60 cm (December, 1993—) was used. The laser pulse repetition rate was 10 Hz. The signals for 2000 laser shots were accumulated to obtain one profile every 4 minutes. Usually, two adjacent profiles were averaged in order to increase the signal to noise ratio, which made the time resolution 8 minutes.

Another Na lidar was installed at the Shigaraki MU observatory in September 1995 by Shinshu University. The basic system of this lidar was the same that used by Nomura *et al.* (1987), with a similar system design to the TMU Na lidar, except that the dye laser was excited by a coaxial flashlamp. As for the receiving system, a Shumidt-Cassegrainian telescope with a diameter of 35.5 cm was used. The laser pulse repetition rates were 0.2 Hz (–October, 1996) and 0.5 Hz (October, 1996–), respectively. The number of accumulations to obtain one profile were 100 (–August, 1996) and 300 (August, 1996–), respectively, with an effective sampling range of 750 m.

The specifications of these lidars are summarized in Table 2, and the locations are displayed on the map in Fig. 1. The two observatories, Shigaraki and Hachioji, are 310 km

Table 2. Specifications of the two sodium density lidars. The Hachioji Na lidar is operated by Tokyo Metropolitan University (TMU), while the Shigaraki Na lidar is operated by Shinshu University.

Hachioji	Shigaraki
(Tokyo Metropolitan Univ.)	(MU Observatory)
35.6°N, 139.4°E	34.9°N, 136.1°E
Nd:YAG/SHG	Flashlamp
pumped DYE	pumped DYE
2.5 pm	3.0 pm
	10.0 pm (1996 Aug)
100 mJ	80–100 mJ
10 Hz	0.2 Hz
	0.5 Hz (1996 Oct)
6 ns	500 ns
0.2 mrad	0.1 mrad
28 cm	35 cm
35 cm (1992 Oct)	
60 cm (1993 Dec)	
0.7 mrad	$1 \sim 3 \text{ mrad}$
3.5 nm (FWHM)	1.0 nm (FWHM)
Photon counting	Photon counting
100 m	750 m
8 minutes	10 minutes
4000 shots	100 shots
	300 shots (1996 Oct)
	(Tokyo Metropolitan Univ.) 35.6°N, 139.4°E Nd:YAG/SHG pumped DYE 2.5 pm 100 mJ 10 Hz 6 ns 0.2 mrad 28 cm 35 cm (1992 Oct.–) 60 cm (1993 Dec.–) 0.7 mrad 3.5 nm (FWHM) Photon counting 100 m 8 minutes

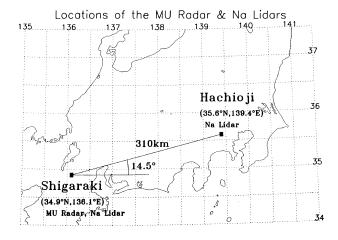


Fig. 1. Locations of the observatories and observation equipment. At Shigaraki, the MU radar and a sodium density lidar are operated by Radio Atmospheric Science Center (RASC), Kyoto University, and Shinshu University, respectively. At Hachioji, a sodium density lidar is operated by Tokyo Metropolitan University (TMU).

apart, and the baseline between the sites is approximately zonally aligned.

2.3 Overview of simultaneous observations

Simultaneous observations of the mesopause region have been carried out with the MU radar and TMU Na lidar since August, 1992, and with the Shigaraki Na lidar since November, 1995. The numbers of simultaneous observations are summarized in Table 3.

Meteor observations with the MU radar are performed on a campaign basis, each observation being usually carried out for a few to ten days. During meteor shower periods in August and November, annual observations have been regularly carried out, while a number of campaigns with optical instruments such as lidars, airglow photometers and imagers have been carried out recently since 1995.

During the MU radar meteor mode observations, Na lidars observed Na density profiles at Hachioji and Shigaraki, although these optical observations were sometimes prevented by cloudy weather conditions, which resulted in a smaller total number of observation days. Between August 1992 and

Table 3. Numbers of observations in a month for the MU radar and sodium lidars. The former correspond to the nights of meteor mode observations, and the latter to the nights of simultaneous observations with the MU radar.

Year Month MU Radar Obs. (Meteor mode) Na lidar, simultaneously 1992 AUG 9 1 — — 1993 JAN 4 2 — — APR 4 2 — — AUG 3 2 — — 1994 NOV 8 2 — — 1995 JUL 6 3 — — 1995 JUL 6 3 — — NOV 9 4 3 1 — — NOV 9 4 3 1 — — 1996 APR 2 1 0 0 0 AUG 5 1 3 0 0 OCT 4 2 3 1 NOV 7 3 4 3 1 DEC 4 2 0 0 0 0<						
(nights) (nights) (nights) (nights)			MU Radar Obs.	Na lid	lar, simultane	ously
1992 AUG 9 1 — — 1993 JAN 4 2 — — APR 4 2 — — AUG 3 2 — — 1994 NOV 8 2 — — 1995 JUL 6 3 — — 1995 JUL 6 3 — — AUG 3 1 — — SEP 5 2 — — NOV 9 4 3 1 1996 APR 2 1 0 0 AUG 5 1 3 0 OCT 4 2 3 1 NOV 7 3 4 3 DEC 4 2 0 0 1997 JAN 10 6 7 4	Year	Month	(Meteor mode)	Hachioji	Shigaraki	(Both)
1993 JAN			(nights)	(nights)	(nights)	(nights)
APR AUG 3 2 — — DEC 3 2 — — 1994 NOV 8 2 — — 1995 JUL 6 3 — — AUG 3 1 — — SEP 5 2 — NOV 9 4 3 1 DEC 5 2 3 1 1996 APR 2 1 1996 APR 2 1 0 0 0 0 0 0 0 0 0 0 0 1 1 1 1 1 1 1	1992	AUG	9	1	_	_
AUG 3 2 — — DEC 3 2 — — 1994 NOV 8 2 — — 1995 JUL 6 3 — — AUG 3 1 — — SEP 5 2 — — NOV 9 4 3 1 DEC 5 2 3 1 1996 APR 2 1 0 0 AUG 5 1 3 0 OCT 4 2 3 1 NOV 7 3 4 3 DEC 4 2 0 0 1997 JAN 10 6 7 4	1993	JAN	4	2	_	_
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NOV 9 4 3 1 DEC 5 2 3 1 1996 APR 2 1 0 0 AUG 5 1 3 0 OCT 4 2 3 1 NOV 7 3 4 3 DEC 4 2 0 0 1997 JAN 10 6 7 4		AUG	3	1	_	_
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1996 APR 2 1 0 0 AUG 5 1 3 0 OCT 4 2 3 1 NOV 7 3 4 3 DEC 4 2 0 0 1997 JAN 10 6 7 4		NOV	9	4	3	1
AUG 5 1 3 0 OCT 4 2 3 1 NOV 7 3 4 3 DEC 4 2 0 0 1997 JAN 10 6 7 4		DEC	5	2	3	1
OCT 4 2 3 1 NOV 7 3 4 3 DEC 4 2 0 0 1997 JAN 10 6 7 4	1996	APR	2	1	0	0
NOV 7 3 4 3 DEC 4 2 0 0 1997 JAN 10 6 7 4		AUG	5	1	3	0
DEC 4 2 0 0 1997 JAN 10 6 7 4		OCT	4	2	3	1
1997 JAN 10 6 7 4		NOV	7	3	4	3
		DEC	4	2	0	0
Total 91 38 23 10	1997	JAN	10	6	7	4
	To	otal	91	38	23	10

January 1997, simultaneous observations with the MU radar were successfully carried out on 38 and 23 nights with the Hachioji and Shigaraki Na lidars, respectively. Simultaneous observations with all three (the MU radar, and two lidars) have been made on 10 nights.

3. Data Analysis

3.1 Analysis of dynamical parameters with the MU radar data

From meteor echo observations with the MU radar, the horizontal wind velocity and diffusion coefficient are obtained with time and height resolutions of 30 minutes and 1 km, respectively, between 75 and 100 km (Nakamura *et al.*, 1997). The relative temperature fluctuation is also estimated from the diffusion coefficient by applying a Boussinesq approximation (Tsutsumi *et al.*, 1996). In this study, we further derive the vertical gradients of the obtained parameters, such as the wind shear and vertical temperature gradients, and then the indices of the atmospheric stability such as The Brunt-Väisälä frequency, N, and Richardson number, Ri, are also calculated.

We estimate horizontal winds, u(z,t) or v(z,t), at altitude z(km) with a height interval of 1 km. Therefore, we can calculate wind shear by taking the difference between the horizontal (zonal or meridional) wind velocities of two neighboring heights in the profile. However, the differentiation shown here will increase the error in wind shear estimation. In order to avoid this, the results are smoothed to

have an effective time and height resolution of $2 \text{ km} \times 1$ hour. In addition to the wind shear of zonal and meridional wind velocities, the total wind shear,

$$\sqrt{\left(\frac{du}{dz}\right)^2 + \left(\frac{dv}{dz}\right)^2} \tag{1}$$

is also calculated with the same time and height resolutions.

The temporal fluctuation of the relative temperature, T'/T_0 ($_0$ denotes the time average), is determined from the ambipolar diffusion coefficient, D, obtained by means of the decay time constant of radar meteor echoes. However, in order to estimate the vertical temperature gradient, which is significantly perturbed by the atmospheric wave activity, we have to determine absolute temperature, T, for each time and height bin in the meteor observations. In this study, we assume that the profile of time-averaged temperature, T_0 , is identical to the model description in CIRA 1986 (COSPAR International Reference Atmosphere), and then the absolute temperature, T, and fluctuation, T', are estimated in each time-height bin of 30 minutes \times 1 km.

From the estimated temperatures, we also calculate the Brunt-Väisälä frequency squared, N^2 , which is expressed as

$$N^2 = \frac{g_0}{T} \left(\frac{dT}{dz} + \frac{g_0}{C_p} \right) \tag{2}$$

where g_0 and C_p are gravitational acceleration and specific heat at constant pressure, respectively. N^2 is also estimated with a time-height resolution of 1 hour \times 2 km, in order to reduce the estimation error. Then, the Richardson number, Ri, which indicates atmospheric stability by wind shears, can be calculated as follows,

$$Ri = \frac{N^2}{\left(\frac{du}{dz}\right)^2 + \left(\frac{dv}{dz}\right)^2}.$$
 (3)

The histograms in Fig. 2 present the distributions of zonal (a), meridional (b) and total (c) shears for the altitude range of all data, i.e., between 75 km and 110 km in August 1995. Zonal and meridional shears can be approximated by a Gaussian distribution with a mean value close to zero and standard deviation of each wind shear of 7.0 m/s/km and 7.8 m/s/km, respectively, while the total shear shows a distribution similar to a χ^2 -distribution with a mean value of 9.0 m/s/km.

The distribution of temperature fluctuation, T', is plotted in Fig. 2(d). The standard deviation of T' (i.e., r.m.s. amplitude) is 15 K. Although the T' distribution can be approximated well by the Gaussian distribution, negative T' values appear slightly more frequently. The distribution of the Brunt-Väisälä frequency squared, N^2 , is plotted in Fig. 2(e) with a Gaussian function as dotted lines. The shaded area indicates the region of convective instability. N^2 also seems to be approximated by a Gaussian distribution as T'. The mean value of N^2 in this month is 4.9×10^{-4} (rad/sec)², corresponding to 4.7 minutes in oscillation period. The standard deviation is 3.1×10^{-4} , indicating that at most times and heights, N is distributed between 3.7 and 7.7 minutes.

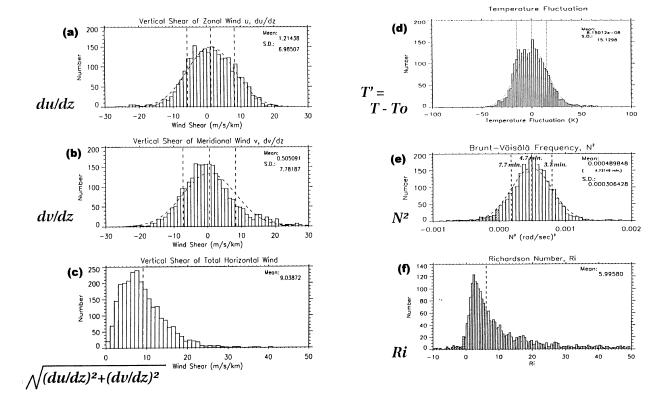


Fig. 2. Distributions of estimated atmospheric parameters observed with the MU radar. Each histogram presents the distribution of wind shears ((a) zonal (du/dz), (b) meridional (dv/dz) and (c) total), (d) temperature fluctuation, T', (e) Brunt-Väisälä frequency squared, N^2 , and (f) Richardson number, Ri, in August 1995 in the altitude range of 75–100 km.

Such statistics are obtained every month and used to determine whether the dynamical parameters at Na_s events are common or significantly strong (or weak).

It should be noted that the MU radar observatory at Shigaraki and the TMU lidar observatory at Hachioji are located about 310 km apart. Thus, the atmospheric parameters observed at Shigaraki are not always the same as those above Hachioji, precisely. However, the atmospheric parameters observed with the MU radar are average values over 200 km in horizontal extent, and are affected mainly by inertia gravity waves with horizontal wavelengths of larger than 500 km, global scale tides, planetary waves and mean winds. Hodograph analyses of the inertia gravity waves seen here revealed that the phase difference of these inertia gravity waves at Shigaraki at the times of the Na_s events shown here are mainly propagated in the meridional direction rather than in the zonal direction, and hence the phase difference between Shigaraki and Hachioji is usually as small as within 30 min-1 hours. It is also noteworthy that the wind and temperature caused by atmospheric tides, which have a large amplitude at this altitude, exhibit a phase difference of only about 13 minutes. Thus, in this study we consider that the background atmospheric parameters for Shigaraki and Hachioji are basically the same with no time lag.

3.2 Analysis of sodium density structure

Although many studies on Na_s have been performed since Clemesha *et al.* (1978) reported Na_s events for the first time at São Paulo, Brazil, the definition of Na_s is not clear, and is sometimes subjective. In order to avoid this uncertainty, we

first describe a more objective method to select Na_s event, and applied to the sodium density data in this study. The criteria are:

- 1) To find an enhancement of Na density as a candidate Na_s event.
- To calculate the full width of half maximum (FWHM) thickness of the layer for each enhancement event, and further select candidates with a FWHM thickness of smaller than 5 km.
- 3) To calculate the ratio of the maximum density at the enhancement to the normal density at the same altitude. The normal density is estimated with the following algorithm, which is the same one as that used by Nagasawa and Abo (1995). The concept is illustrated in Fig. 3.
 - (a) To determine the peak density, ρ_p , and the corresponding height, H_p , of a candidate Na_s.
 - (b) To detect the upper and lower heights, H_{uh} and H_{lh} , where Na density becomes half the maximum, i.e., $\rho_p/2$. If there are more than two H_{uh} or H_{lh} , choose the nearest one to H_p .
 - (c) Estimate the altitudes of the upper and lower ends, H_{ue} and H_{le} , of the candidate as follows

$$H_{ue} = H_p + 2(H_{uh} - H_p),$$

 $H_{le} = H_p - 2(H_p - H_{lh}).$

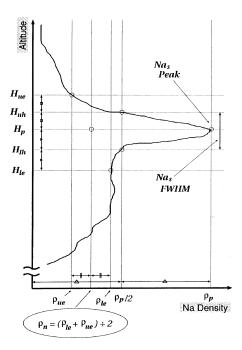


Fig. 3. The concept of FWHM of Na_s and estimation of the density of the normal Na layer (see text for details).

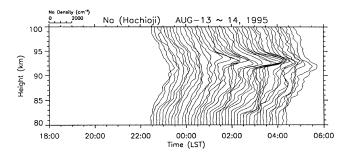


Fig. 4. An example of Na density profiles observed with the Hachioji Na lidar on August 13/14, 1995.

- (d) To determine the Na density, ρ_{ue} and ρ_{le} , at H_{ue} and H_{le} .
- (e) The density of the normal Na layer, ρ_n , at H_p is calculated as

$$\rho_n = \frac{\rho_{ue} + \rho_{le}}{2}.$$

The ratio of the maximum Na density to that of the normal layer is defined as ρ_p/ρ_n . If this ratio is greater than two, the event is judged to be an Na_s event.

Figure 4 shows the profiles of an Na_s event detected by applying the above procedure, on August 13/14, 1995, at Hachioji. During this night the background structure of the sodium layer profile did not change significantly, and fluctuations with downward or upward phase propagations associated with atmospheric waves were not clearly observed. Starting at 0:30 LST (= JST = UT + 9), a significant Na_s event appeared at 93 km, and continued for about 3 hours,

with the peak moving downward slightly for about one hour, and then slowly upward at a speed of about 1 km/hour. At 1:57 LST, the peak Na_s density reached a maximum at 92.8 km, the ratio of the Na_s maximum to the normal layer being 2.8.

4. Results

4.1 Na_s events

Na_s events are selected according to the criteria described in the previous section, and eleven and four Na_s events were detected at Hachioji and Shigaraki, respectively, during the simultaneous observation period. The characteristics of the Na_s events are summarized in Table 4, and also in Fig. 5.

Figure 5(a) shows the distribution of $\mathrm{Na_s}$ duration. Most of the events continued for a few tens of minutes to one hour. However, three events at Hachioji and one at Shigaraki continued for more than two hours. The mean value and standard deviation of the $\mathrm{Na_s}$ duration are 74 and 46 minutes, respectively. Nagasawa and Abo (1995) reported the statistics for $\mathrm{Na_s}$ based on observations with the same lidar system but for a larger database of 37 $\mathrm{Na_s}$ events observed in 1991–1993 at Hachioji, and showed that the mean value and standard deviation of the $\mathrm{Na_s}$ duration are 110 minutes and 63 minutes, respectively, which are slightly larger but not significantly different from our results.

The distribution of Na_s altitude at the maximum Na density is plotted in Fig. 5(b). Na_s events seem to be classifiable into two types according to the peak altitude; i.e. between about 91 and 97 km altitude with a mean at about 94 km, and above 99 km.

The distribution of FWHM of the Na_s layer, shown in Fig. 5(c), indicates that a thickness of 3 to 4 km is commonly seen, although it should be noted that with our criteria the maximum FWHM is limited to 5 km. The FWHM of Na_s observed here is slightly broader than in other studies, such as those of Nagasawa and Abo (1995) and Qian *et al.* (1998), in which 1–3 km was typical. It should be noted that Nagasawa and Abo (1995) calculated the FWHM only for Na_s density after subtracting the normal Na layer, however, in the current analysis the FWHM is determined for the total Na density, and therefore this difference in the analysis could cause the FWHM values in Fig. 5 to be slightly larger than those of Nagasawa and Abo (1995).

The ratio of the peak Na density at Na_s to the normal density of the Na layer is plotted in Fig. 5(d), which shows 3–4 is typical, and that is sometimes becomes larger than 5. This tendency is similar to the previous observations at Hachioji (Nagasawa and Abo, 1995) and Hawaii in the ALOHA/ANLC-93 campaign (Qian *et al.*, 1998).

Thus, in general the observed characteristics of the Na_s layer shown in Table 4 and Fig. 5 are similar to those reported by Nagasawa and Abo, (1995), which suggests our criteria for Na_s recognition shown in the previous section work reasonably well.

Nagasawa and Abo (1995) indicated a strong correlation of Na_s with Es occurrences. The existence of Es around the time of each Na_s event observed at Shigaraki and Hachioji by means of ionosondes is also displayed in Table 4. Eight of the 15 events were accompanied by Es, mainly in the summer, with one event in January.

 $Table \ 4. \ \ Detected \ Na_s \ events, and summary \ of the \ results \ of \ correlation \ analysis \ between \ Na_s \ occasions \ and \ the \ dynamical \ structures \ of \ the \ atmosphere.$

Hachioji (35.6°N, 139.4°E)

Observation	Na _s	Duration		At the peak density					
date	period		Time	Height	FWHM	Density	Density	Es	
(DMY)	(LST)	(minutes)	(LST)	(km)	(km)	(10^3 cm^{-3})	ratio		
06/07-AUG-1992	19:56-20:22	26	20:09	95.5	3.5	2.2	3.7	Y	
06/07-AUG-1992	21:33-21:57	24	21:39	99.3	3.5	1.7	3.3	N	
28/29-JAN-1993	23:44-0:47	63	00:03	93.9	1.9	3.0	2.7	N	
24/25-JUL-1995	20:14-20:38	16	20:26	91.9	1.9	2.7	3.3	Y	
24/25-JUL-1995	22:38-23:25	47	23:22	96.5	3.4	2.0	3.5	Y	
26/27-JUL-1995	19:46-20:51	65	20:12	95.5	3.0	2.5	3.7	Y	
26/27-JUL-1995	23:29-0:16	47	23:41	93.9	1.6	3.9	5.4	Y	
13/14-AUG-1995	0:18-2:49	151	01:57	92.8	3.1	4.4	2.8	N	
20/21-SEP-1995	19:57-21:14	77	20:27	94.4	2.3	4.7	3.2	N	
22/23-APR-1996	22:04-0:06	122	23:31	93.4	2.4	1.6	3.5	N	
11/12-JAN-1997	23:05-1:10	127	23:31	96.9	1.8	4.9	4.3	Y	

Shigaraki (34.9°N, 136.1°E)

Observation	Na _s	Duration		At the peak density				
date	period		Time	Height	FWHM	Density	Density	Es
(DMY)	(LST)	(minutes)	(LST)	(km)	(km)	(10^3 cm^{-3})	ratio	
23/24-DEC-1995	22:18–22:45	27	22:18	92.6	3.8	9.1	3.2	N
11/12-AUG-1996	21:28-22:13	45	21:55	99.4	3.0	8.0	3.2	Y
11/12-AUG-1996	23:12-1:24	132	0:31	93.4	4.5	10.5	5.6	Y
15/16-NOV-1996	21:56-22:17	21	21:56	91.1	3.8	7.3	3.5	N

	Vertical shear of horizontal wind			Tempe	erature	Brunt-Väisälä	Richardson
Zonal	Meridional	Total	Direction	T'	T	Frequency, N^2	Number, Ri
(m/s/km)	(m/s/km)	(m/s/km)	(Azimuth, $^{\circ}$)	(K)	(K)	$(\times 10^{-4} \text{ rad/sec})^2$	
+9.2	-3.9	10.0	+113	-10	169	3.21	3.58
+5.2	-14.8	15.7	+161	_	_	_	_
-5.3	-20.9	21.6	-166	+16	206	2.52	0.54
-17.7	-3.4	18.1	-100	-20	148	5.56	1.70
+2.2	+1.0	2.4	+65	+15	195	5.77	164
+2.7	+2.5	3.7	+48	-1	179	7.03	50.8
-2.9	+9.0	9.4	-18	-20	153	9.96	11.2
+15.6	-16.1	22.4	+136	+29	201	6.13	1.22
-13.6	+5.4	14.7	-69	+2	178	2.34	1.09
-14.8	+18.3	23.5	-39	-19	158	5.43	2.86
+9.6	-10.1	13.9	+137	-23	168	3.02	1.56

	Vertical shear of horizontal wind					Brunt-Väisälä	Richardson
Zonal	Meridional	Total	Direction	T'	T	Frequency, N^2	Number, Ri
(m/s/km)	(m/s/km)	(m/s/km)	(Azimuth, $^{\circ}$)	(K)	(K)	$(\times 10^{-4} \text{ rad/sec})^2$	
-2.8	+14.2	+14.5	-11	-17	176	5.47	14.0
_	_	_	_	_	_	_	_
+1.7	+7.7	+7.9	+13	+8	180	3.67	1.7
-25.5	-3.6	+25.8	-98	-2	188	4.64	23.2

Nas Characteristics

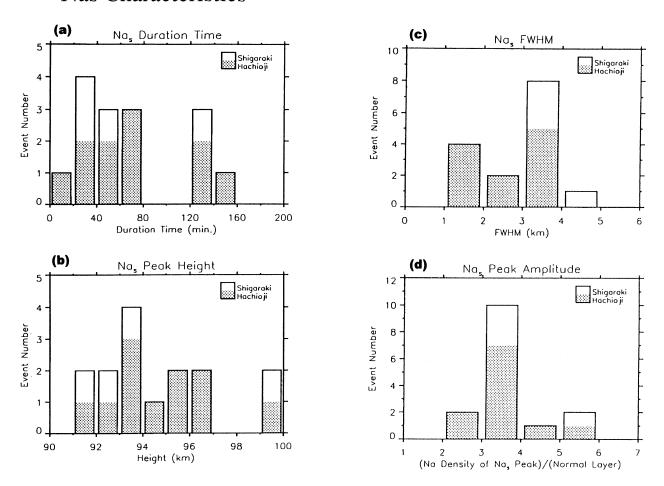


Fig. 5. Distributions of Na_s characteristics. The histograms present (a) Na_s duration time, (b) height, (c) FWHM, and (d) amplitude at the Na_s density maximum. The open and shaded boxes correspond to the events observed at Shigaraki and Hachioji, respectively.

4.2 Simultaneous observations: case studies

The first two panels in Fig. 6 present the profiles and a contour plot of the sodium density between 18:00 LST and 06:00 LST on September 20/21, 1995, observed with the TMU Na lidar at Hachioji. The other panels in the same figure show the zonal and meridional wind velocities, and their vertical shear, temperature fluctuations, Brunt-Väisälä frequencies, and Richardson numbers, observed simultaneously with the MU radar at Shigaraki. During the observation period on this night, a normal Na layer moved downward from 94 km at 19:00 LST to 90 km at 21:00 LST, with a vertical speed of about 2 km/hour. A strong Na_s event started at 19:57 LST at about 95 km altitude and the density reached the maximum at 20:27 LST at 94.4 km altitude, decreasing slowly afterwards. This Na_s event continued at least until 21:14 LST when the lidar observation was interrupted, with a downward motion with a velocity of about 1 km/hour. It is noteworthy that the downward velocity of the Na_s was smaller than that of the normal Na layer. During the Na_s event, strong westward shear of the zonal wind was observed between 19:30 and 21:00 LST and between 94 and 95 km altitude, which is about the time and altitude of the Na_s occurrence. The ionosondes at both Shigaraki and Kokubunji (near Hachioji)

did not detect a sporadic E layer around the time of this event.

The wind velocities, u and v, showed strong oscillation with a period of around 12 hours, especially above 90 km, which is probably a semidiurnal tide or an inertia gravity wave. The peak amplitude reached 60 m/s, and the downward phase progression suggests an upward energy propagating wave. Accordingly, the wind shears, du/dz and dv/dz, showed similar downward phase progression, but the vertical scale was shorter and exhibited a shorter period component superposed on the 12 hours variation, suggesting multiple waves were propagating at the same time. At 20:30 LST and 94 km altitude, when and where the Na density of the Na_s layer became maximum, the intensity of the zonal wind shear, du/dz, was -13.6 m/s/km. The standard deviation of the zonal shear, i.e., the r.m.s. amplitude observed in this month, was 6.8 m/s/km. Thus, the westward shear observed here was significantly large compared with the statistical fluctuation of the zonal shear. du/dz reached its maximum of -19.6 m/s/km at 20:00 LST and 95 km altitude, which preceded the time of the Na_s density maximum by 30 minutes and 1km above the Na_s layer. On the other hand, dv/dz at the Na_s maximum was as weak as -5.4 m/s/km. In this case, the Na_s event exhibited a strong positive correlation with

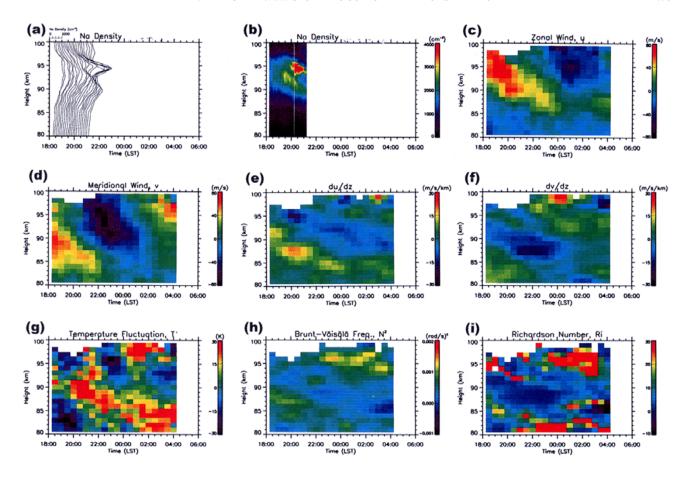


Fig. 6. (a), (b) Na density profiles and a contour plot obtained with the Hachioji Na lidar on September 20/21, 1995. Contour plots of the zonal (c) and meridional (d) wind velocities, zonal (e) and meridional (f) wind shears, temperature fluctuation (g), Brunt-Väisälä frequency squared (h), and Richardson number (i) observed simultaneously with the MU radar at Shigaraki.

the zonal wind shear. This strong shear could be caused by the superpositioning of multiple waves, as discussed above. As can be seen in Fig. 6, the temperature fluctuation, T', at the Na_s peak was negative, and the Brunt-Väisälä frequency squared, N^2 , was not very small. However, the Richardson number, Ri, was as small as 1.1, which corresponds to the strong zonal wind shear.

Another case study ivolving the comparison between the Na_s and radar observations is shown in Fig. 7 for the data obtained at Hachioji on August 13/14, 1995. An Na_s event reached the maximum density at 2:00 LST and 93 km altitude. The horizontal wind, u and v, showed a strong wavelike structure with a period of about 15 hours, with an amplitude of over 50 m/s, which is considered to be an inertia gravity wave with upward energy propagation. The wind shears, du/dz and dv/dz, showed complicated variation, which again suggests superpositioning of multiple waves. However, a significant peak of du/dz and dv/dz was seen at the peak Na_s , although the direction of the shear was different from in the previous case.

On the other hand, around the time and height of the Na_s event, temperature enhancement was observed. At the density maximum of the Na_s event, T' was 29 K, corresponding to the local maximum of T', where the absolute temperature, T, was 201 K. The standard deviation of T' observed in this month was 15 K, so the T' value of 29 K was consid-

erably large in comparison with the standard deviation. In this case, the temperature enhancement and Na_s occurrence were observed simultaneously, which is similar to the results obtained by Gardner *et al.* (1995).

In Fig. 7, the Richardson number, Ri, at the maximum sodium density of Na_s, i.e., at 2:00 LST and 93.0 km altitude, was 1.2. The value of Brunt-Väisälä frequency squared, N^2 , at that time and height was 6.1×10^{-4} (rad/sec)², corresponding to 4.2 minutes in period. N^2 was slightly large compared to the average value of 4.9×10^{-4} (rad/sec)² and the standard deviation of 3.1×10^{-4} (rad/sec)². However, Ri was significantly smaller than the mean Ri in this month, which was 6.0. If we look for a local minimum of Ri, it occurred at 2:00 LST and 94 km altitude, which is 1 km above the Na_s layer, and the corresponding Ri was also 1.2, and Ri did not become negative.

It is notable that the stability indices obtained in this study were for a fairly wide time and height scale with a resolution of 200 km in the horizontal extent and 2 km in thickness, considering the meteor echo distribution and smoothing procedure in the analysis. Thus, the indices here are mainly due to large scale atmospheric waves, such as tides and inertia gravity waves both of which have large amplitudes at this altitude range. Therefore, the superposed small scale gravity waves are likely to make the indices such as N^2 and Ri even smaller, and in the case in Fig. 7, it is probable that the

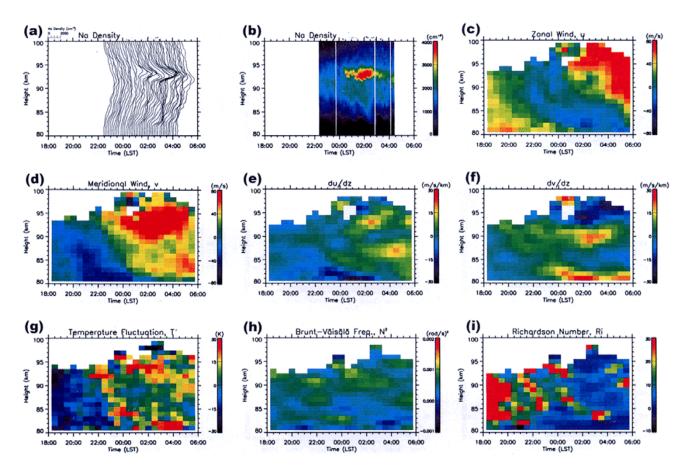


Fig. 7. The same as Fig. 6, except that the data are for August 13/14, 1995.

region of the Na_s peak could be in the region of gravity wave breaking and subsequent heating, as supposed by Zhou *et al.* (1993).

4.3 Simultaneous observation: statistical approach

We perform the same analysis for all the 15 Na_s events observed, in order to discuss the dynamical background of Na_s more statistically, although the number of samples is not so large. The results are summarized in Table 4. The vertical shears of zonal and meridional winds, temperature fluctuations, Brunt-Väisälä frequencies, and Richardson numbers are estimated from the MU radar data at the time and altitude of the Na_s density maximum.

The distribution of zonal wind shears at the Na_s maximum density is plotted in the top panel of Fig. 8(a). The normalized shear, i.e., the wind shear normalized as to the standard deviation (r.m.s. amplitude) of the shear in each month, is plotted in the bottom panel of the same figure.

In Fig. 8(a), for seven of the 14 Na_sevents, i.e., half the zonal shears, du/dz, were shown to be westward, and the rest to be eastward. Similarly, northward and southward shears were evenly observed in Fig. 8(b). Thus, no clear preferential shear direction was found. For both zonal and meridional wind shears four of the 14 events showed more than twice the monthly standard deviation, and the number of cases with shears smaller than the standard deviation was small, as shown in the bottom panels of Fig. 8(a) and (b).

Considering the fact that there were many cases of strong shear of the horizontal wind, we now investigate the magnitud of the total shear of the horizontal wind, as shown in Fig. 8(c). The total wind shear normalized as to the mean amplitude in each month is also plotted in the bottom panel. It is notable that the statistical study on the MU radar wind data mentioned earlier in this paper showed that the occurrence rate of total shear exceeding the monthly mean is about 43%, and only 6.2% of the data exceed twice the mean. On the other hand, for the amplitudes of the total shears of the horizontal wind at the Na_s density maxima, eleven of the 14 events, i.e., 78%, exhibit wind shears exceeding the monthly mean, and five, i.e., 35% of all events, exhibit shears of more than twice the monthly mean. Shears were less than the monthly mean for only three events, which is about 21% of all events. Therefore, the total wind shears around the time and altitude of Na_s layers were significantly large for our observations, which is similar to the finding of Qian et al. (1998).

In order to confirm the significance of the wind shear around the Na_s , the distribution of normalized shears observed at 3 km below the Na_s peak altitude is compared in Fig. 9. The wind shears at 3 km below were generally smaller. Although there were not many cases when the wind shear could be determined by means of radar observation at 3 km higher than Na_s , all three cases showed smaller wind shears than those at Na_s peaks. Therefore, we can conclude that the total shear at the Na_s peaks was considerably large, compared with both the statistical average and other heights.

It is noteworthy that among the seven Na_s events with total

(a) Zonal shear at Nas peak (b) Meridional Shear at Nas peak (c) Total shear at Nas peak (c) Total shear at Nas peak (d) Total shear at Nas peak (e) Total shear at Nas peak (f) Total shear at Nas peak (g) Total s

Fig. 8. Distributions of (a) zonal, (b) meridional and (c) total wind shears at the Na_s density maximum. The top panels show the amplitudes of shear, while the bottom panels present the ratios, i.e., the wind shear normalized as to the standard deviation of the shear (a, b) or the average of the shear (c) in each month

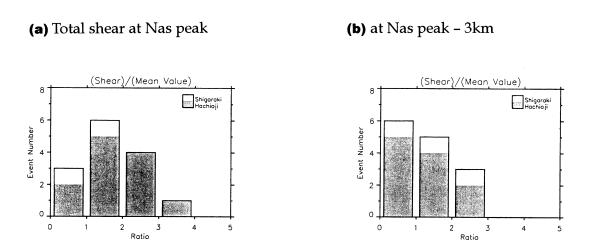


Fig. 9. Distributions of total wind shears at (a) the Na_s density maximum, and (b) 3 km below the Na_s density maximum. Total wind shears are normalized as to the monthly average.

wind shears exceeding 1.6 times the mean value, only one event was accompanied by Es. Therefore, the Na_s events without an Es occurrence were more significantly correlated with strong total wind shears (see Fig. 10).

Next, we investigate the temperature at the times and heights of Na_s events. The results shown in Table 4 are plotted as histograms in Fig. 11, where (a) temperature fluctuation, T', (b) normalized T' as to the standard deviation in each month, and (c) absolute temperature, T, at the Na_s density maximum are displayed. In these figures, eight and five out of 13 events showed negative and positive temperature fluctuation, T', and therefore it is recognized that temperatures are rather lower at the times and heights of Na_s occurrences than the average values. From these results, a correlation of Na_s events with a temperature decrease, instead of enhancement which was proposed in the theory of

Zhou and Mathews (1995), is more probable, but it is difficult to conclude this definitely with the small samples in the current study.

Similarly, the Brunt-Väisälä frequency squared, N^2 , and Richardson number, Ri, are also investigated and plotted as histograms in Fig. 12 together with the N^2 fluctuations, $N^{2'}$, i.e., normalized as to the standard deviation in each month. N^2 at the Na_s density maximum seems to be mainly distributed between $2-6 \times 10^{-4} \, (\text{rad}^2/\text{sec}^2)$ and is close to the average values because the ratio is mainly distributed within \pm 0.5. In comparison with the distribution of the average Ri presented in Fig. 2(f), the Richardson numbers were generally small for most of the Na_s events. This finding is associated with the fact that large wind shears are observed at around the Na_s density maximum. To summarize, the Brunt-Väisälä frequency, N, did not show the tendency of an

Total shear at Nas peak

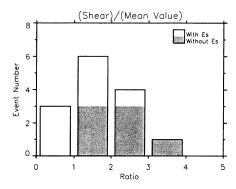
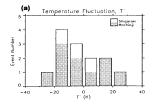
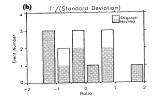


Fig. 10. Distributions of the total wind shear at the Na_s density maximum. The shaded and open areas correspond to Na_s events without and with an Es layer, respectively.

Temperature deviation at Nas peak





Temperature at Nas peak

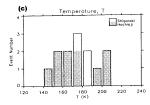
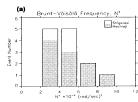
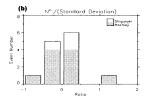


Fig. 11. Distributions of the temperature fluctuation, as presented in Fig. 8(a), (b), and temperature at the Na_s density maximum (c).

Brunt-Väisälä frequency at Nas peak





Richardson number at Nas peak

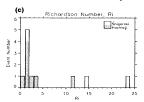


Fig. 12. Distributions of the Brunt-Väisälä frequency squared, N², as presented in Fig. 8(a), (b), and Richardson number at the Na_s density maximum (c).

unstable atmosphere, but the Richardson number, Ri, was comparatively small.

5. Discussion and Concluding Remarks

Gardner et al. (1995) reported an Na_s event accompanied with a strong wind shear observed by a Na wind/temperature lidar at Haleakala, Maui (20.8°N, 156.2°W) during the ALOHA-93 campaign. On October 6, 1993, they observed a strong vertical shear of the zonal wind of 42 m/s/km, which reached the shear instability limit under the assumption that the Brunt-Väisälä frequency corresponds to the period of 5 minutes at the altitude of 1 km below a prominent Na_s layer. Qian et al. (1998) further investigated 43 Na_s events observed during the same campaign and concluded that strong shear of horizontal winds were observed during the majority of the Na_s events and the maximum shears were located on average about 0.5 km below the Na_s layer. In most cases, the meridional shears were significantly stronger than the zonal ones. The maximum total shears were generally directed to the north-northeast and to the south.

On the other hand, temperature variations are also thought to cause Na abundance variation through various chemical processes. Zhou $et\ al.$ (1993) argued that a temperature fluctuation induced by gravity waves may cause Na_s layer formation since the production rate of Na is known to be a sensitive function of temperature. The local enhancement of temperature caused by gravity wave breaking is thought to accelerate the chemical reaction generating atomic sodium from Na compounds. Gardner $et\ al.$ (1995) observed two Na_s events associated with temperature enhancement, which was more than 40 K, on their sodium lidar observations, but on further analysis using the same dataset, Qian $et\ al.$ (1998) found that the temperature enhancement was not always associated with Na_s layers.

In this study, the correlations between the sporadic sodium layer at mid-latitude and dynamical parameters have been investigated by means of the MU radar and the two sodium lidars at Shigaraki and Hachioji. Because of the observational height range of the MU radar, comparisons were limited to the altitudes below 100 km. Wind shears, temperature variations, and atmospheric stabilities indices have been investigated at the times and heights of occurrence of Na_s layers detected by the sodium lidars.

In the current study, we have shown the case of an Na_s event with a strong westward wind shear, and another case with a significant temperature increase. In both cases Ri was fairly small, as discussed in the previous section. However, in a statistical sense, by surveying all fifteen Na_s events observed here, we can conclude that a strong wind shear is more likely to accompany the Na_s occurrence. Especially when sporadic E layers were not existent, such shear was quite large. However, no other clear correlation with other dynamical parameters was seen in the comparisons. The results shown here are very similar to the report by Qian et al. (1998) on the lidar measurements during the ALOHA-93 campaign in Hawaii ($\sim 20^{\circ}$ N), and therefore it could be possible that the Na_s in Hawaii and Japan, both located at a comparatively low latitude, are caused by a similar generation mechanism, although the details of the mechanism are not clear at the moment. We should also note that propagating diurnal tides (positive modes) in the horizontal winds are significantly strong at the latitudes of Hawaii and Japan (20–35 deg), and therefore the contribution of tides to the wind shears could be large, as well as inertia gravity waves.

In this study we have obtained some characteristics of Na_s and associated dynamical characteristics of mid-latitude Na_s layers. Further comparisons of radar and lidar measurements, with better time-height resolutions and preferably at the same location (in order to avoid the uncertainty of the spatial differences of dynamical parameters observed with the radar) in the future are desirable to clarify the generation mechanism of Na_s .

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References

- Batista, P. P., B. R. Clemesha, and D. M. Simonich, Characteristics of the sporadic sodium layers observed at 23°S, J. Geophys. Res., 94, 15349– 15358, 1989.
- Beatty, T. J., R. L. Collins, C. S. Gardner, C. A. Hostetler, C. F. Sechrist, Jr., and C. A. Tepley, Simultaneous radar and lidar observations of sporadic E and Na layers at Arecibo, *Geophys. Res. Lett.*, **16**, 1019–1022, 1989.
- Clemesha, B. R., V. W. J. H. Kirchhoff, D. M. Simonich, and H. Takahashi, Evidence of an extra-terrestrial source for the mesospheric sodium layer, *Geophys. Res. Lett.*, 5(10), 873–876, 1978.
- Clemesha, B. R., V. W. J. H. Kirchhoff, D. M. Simonich, H. Takahashi, and P. P. Batista, Spaced lidar and nightglow observations of an atmospheric sodium enhancement, *J. Geophys. Res.*, 85, 3480–3484, 1980.
- Clemesha, B. R., P. P. Batista, and D. M. Simonich, Concerning the origin of enhanced sodium layers, *Geophys. Res. Lett.*, 15, 1267–1270, 1988.
- Clemesha, B. R., P. P. Batista, and D. M. Simonich, Formation of sporadic sodium layers, J. Geophys. Res., 101, 19701–19706, 1996.
- Cox, R. M. and J. M. C. Plane, An ion-modecule mechanism for the formation of neutral sporadic Na layers, *J. Geophys. Res.*, 103, 6349–6359, 1998.
- Fukao, S., T. Sato, T. Tsuda, S. Kato, K. Wakasugi, and T. Makihira, The MU Radar with an active phased array system: 1. Antenna and power amplifiers, *Radio Sci.*, 20, 1155–1168, 1985a.
- Fukao, S., T. Tsuda, T. Sato, S. Kato, K. Wakasugi, and T. Makihira, The MU Radar with an active phased array system: 2. In-house equipment, *Radio Sci.*, 20, 1169–1176, 1985b.
- Gardner, C. S., D. G. Voelz, C. F. Sechrist, Jr., and A. C. Segal, Lidar studies of the nighttime sodium layer over Urbana, Illinois, 1, Seasonal and nocturnal variations, *J. Geophys. Res.*, **91**, 13659–13673, 1986.
- Gardner, C. S., D. C. Senft, and K. H. Kwon, Lidar observations of a substantial sodium depletion in the summertime Arctic mesosphere, *Nature*, 332, 142–144, 1988.
- Gardner, C. S., X. Tao, and G. C. Papen, Observations of strong wind shears and temperature enhancement during several sporadic Na layer events above Haleakala, *Geophys. Res. Lett.*, **22**(20), 2809–2812, 1995.
- Granier, C. and G. Megie, Daytime lidar measurements of the mesospheric sodium layer, *Planet. Space Sci.*, **30**, 169–177, 1982.

- Hansen, G. and U. von Zahn, Sudden sodium layers in polar latitudes, J. Atmos. Terr. Phys., 52, 585–608, 1990.
- Kane, T. J., C. S. Gardner, Q. Zhou, J. D. Mathew, and C. A. Tepley, Lidar, radar and airglow observations of a prominent sporadic Na/sporadic E layer event at Arecibo during AIDA-89, J. Atmos. Terr. Phys., 55(3), 499–511, 1993.
- Kirkwood, S. and U. von Zahn, On the role of auroral electric fields in the formation of low altitude sporadic-E and sudden sodium layers, *J. Atmos. Terr. Phys.*, 53, 389–407, 1991.
- Kobayashi, K., T. Kitahara, T. D. Kawahara, Y. Saito, A. Nomura, T. Nakamura, T. Tsuda, M. Abo, C. Nagasawa, and M. Tsutsumi, Simultaneous measurements of dynamical structure in the mesopause region with lidars and MU radar, *Earth Planets Space*, 51, this issue, 731–739, 1999
- Kwon, K. H., D. C. Senft, and C. S. Gardner, Lidar observations of sporadic sodium layers at Mauna Kea Observatory, Hawaii, *J. Geophys. Res.*, 93, 14,199–14,208, 1988.
- Nagasawa, C. and M. Abo, Lidar observations of a lot of sporadic sodium layers in mid-latitude, Geophys. Res. Lett., 22(3), 263–266, 1995.
- Nakamura, T., T. Tsuda, M. Tsutsumi, K. Kita, T. Uehara, S. Kato, and S. Fukao, Meteor wind observations with the MU radar, *Radio Sci.*, 26, 857–869, 1991.
- Nakamura, T., T. Tsuda, S. Fukao, H. Takahashi, P. P. Batista, R. A. Buriti, M. Tsutsumi, M. Ishii, K. Igarashi, H. Fukunishi, Y. Yamada, A. Nomura, T. D. Kawahara, K. Kobayashi, C. Nagasawa, M. Abo, and M. J. Taylor, Studies of the MLT regions using the MU radar and simultaneous observations with OH spectrometer and other optical instruments, *Adv. Space Res.*, 19, 643–652, 1997.
- Namboothiri, S. P., T. Tsuda, M. Tsutsumi, T. Nakamura, C. Nagasawa, and M. Abo, Simultaneous observations of mesospheric gravity waves with the MU radar and a sodium lidar, *J. Geophys. Res.*, 101, 4057–4063, 1996.
- Nomura, A., T. Kano, Y. Iwasaka, H. Fukunishi, T. Hirasawa, and S. Kawaguchi, Lidar observations of the mesospheric sodium layer at Syowa station, Antarctica, *Geophys. Res. Lett.*, 14, 700–703, 1987.
- Qian, J., Y. Gu, and C. S. Gardner, Characteristics of the sporadic Na layers observed during the ALOHA/ANLC-93 campaigns, *J. Geophys. Res.*, **103**(D6), 6333–6347, 1998.
- Tsutsumi, M., T. Tsuda, T. Nakamura, and S. Fukao, Wind velocity and temperature fluctuations due to a 2-day wave observed with radio meteor echoes, *J. Geophys. Res.*, **101**(D5), 9425–9432, 1996.
- von Zahn, U., P. von der Gathen, and G. Hansen, Forced release of sodium from upper atmospheric dust particles, *Geophys. Res. Lett.*, **14**, 76–79, 1987
- von Zahn, U. and T. L. Hansen, Sudden neutral sodium layers: a strong link to sporadic E layers, J. Atmos. Terr. Phys., 50, 93–104, 1988.
- von Zahn, U. and E. Murad, NaHCO3, a source of Na atoms for sudden sodium layers, Geophys. Res. Lett., 17, 147–150, 1990.
- Zhou, Q. and J. D. Mathews, Generation of sporadic sodium layers via turbulent heating of the atmosphere?, *J. Atmos. Terr. Phys.*, **57**(11), 1309–1319, 1995.
- Zhou, Q., J. D. Mathews, and C. A. Tepley, A proposed temperature dependent mechanism for the formation of sporadic sodium layers, *J. Atmos. Terr. Phys.*, 55(3), 513–521, 1993.

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